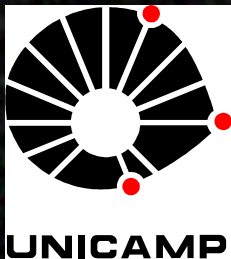


Desvendando o mundo das biomoléculas através de técnicas avançadas baseadas em fontes de raio-X de alta intensidade

Arnaldo Naves de Brito



Departamento de Física Aplicada-IFGW
Universidade Estadual de Campinas – UNICAMP
13083-859 Campinas- Sao Paulo

Group and collaborations



1. *Universidade de Brasília, Instituto de Física- Brasília DF*
2. *Laboratório Nacional de Luz Síncrotron (LNLS) Campinas-SP*
3. *Universidade Federal do ABC, Santo André-SP*
4. *Universidade Federal da Bahia, Barreiras-BA*

More Colaborations



Geraldo Gerson B. de Souza
UFRJ-Rio de Janeiro-RJ



Ênio Frota da Silveira
Lúcio Farenzena
Peter Iza
Ana Mônica
PUC-Rio de Janeiro-RJ

Claudia Lage
Douglas Galante
Ivan Soares
Instituto de Biofisica Carlos Chagas Filho - UFRJ

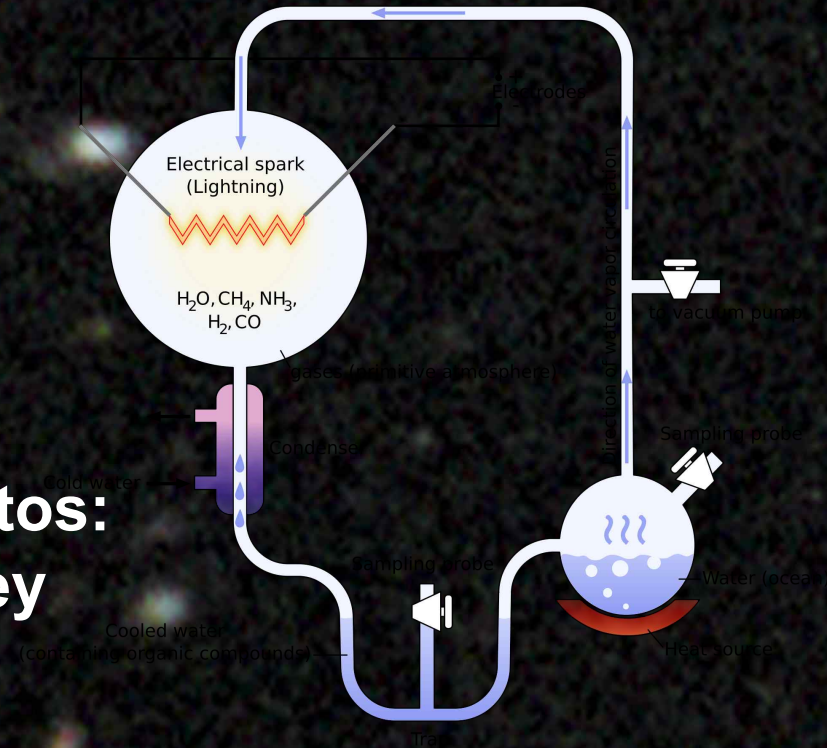


Cenários da evolução pré-biótica

A comunidade científica concorda em geral: Os aminoácidos, e bases nitrogenadas (nucleotídeos e nucleocídeos) foram os precursores formando peptídeos, proteínas e os polímeros RNA e DNA.

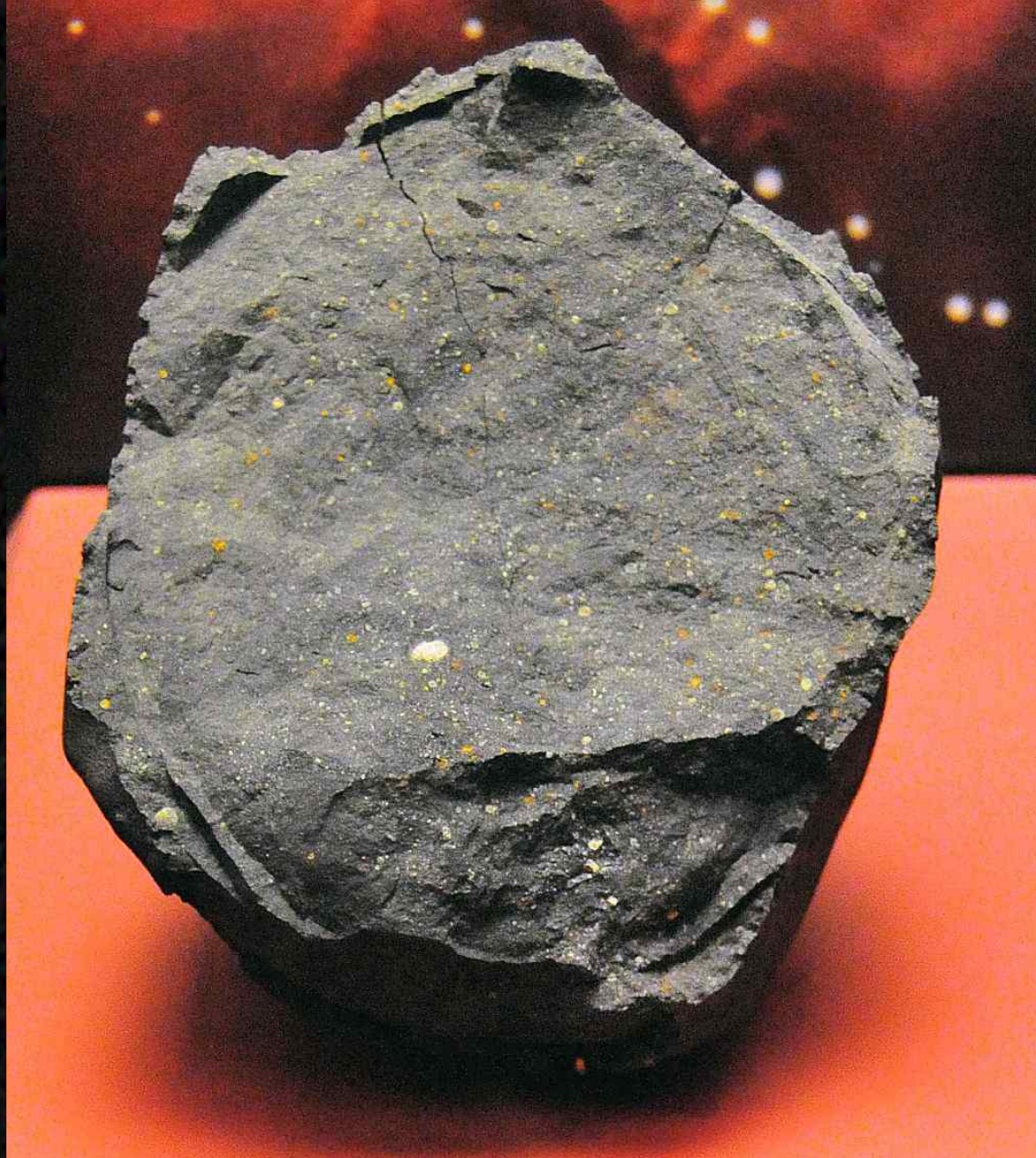


**Experimentos:
Miller e Urey
1953
Crédito: NASA**



Em 1969 um meteoro
“aterrisou” em Murchison,
Austrália. Foi mostrado que
o mesmo continha os
mesmos compostos
orgânicos e amoniácidos e
nas MESMAS proporções
daquelas originadas no
experimento de Miller!

Compound class ^[9]	Concentration (ppm)
Amino acids	17-60
Aliphatic hydrocarbons	>35
Aromatic hydrocarbons	3319
Fullerenes	>100
Carboxylic acids	>300
Hydrocarboxylic acids	15
Purines and Pyrimidines	1.3
Alcohols	11
Sulphonic acids	68
Phosphonic acids	2

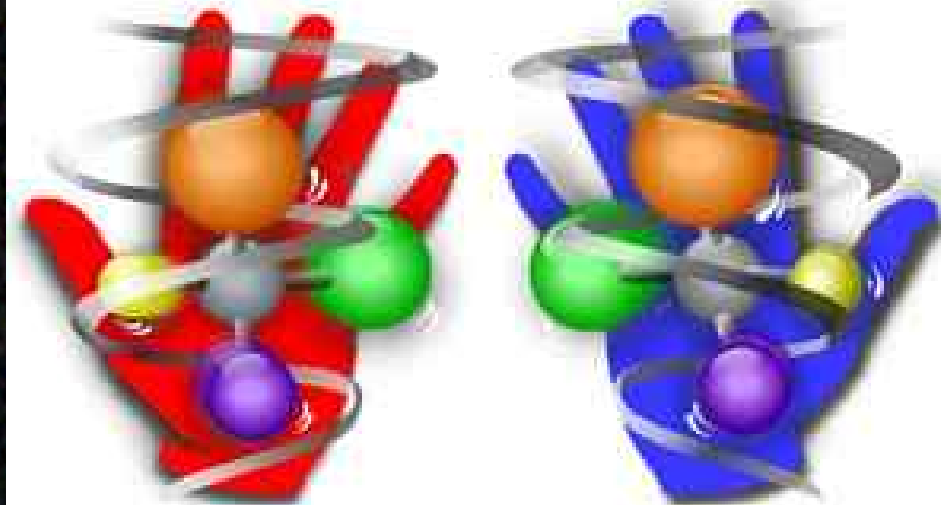
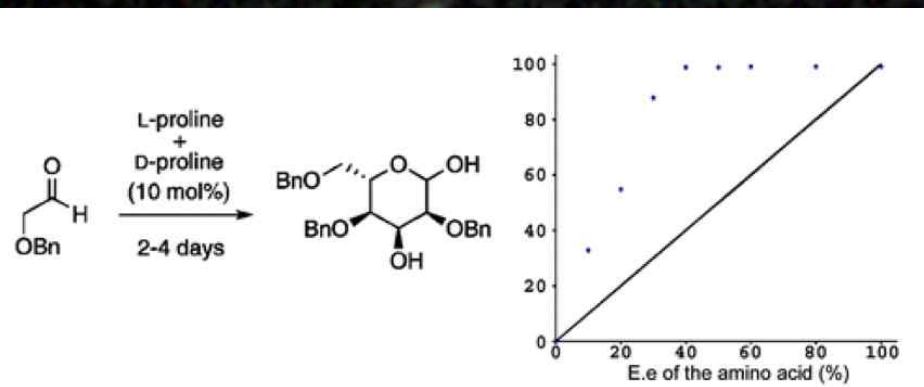


Problema:

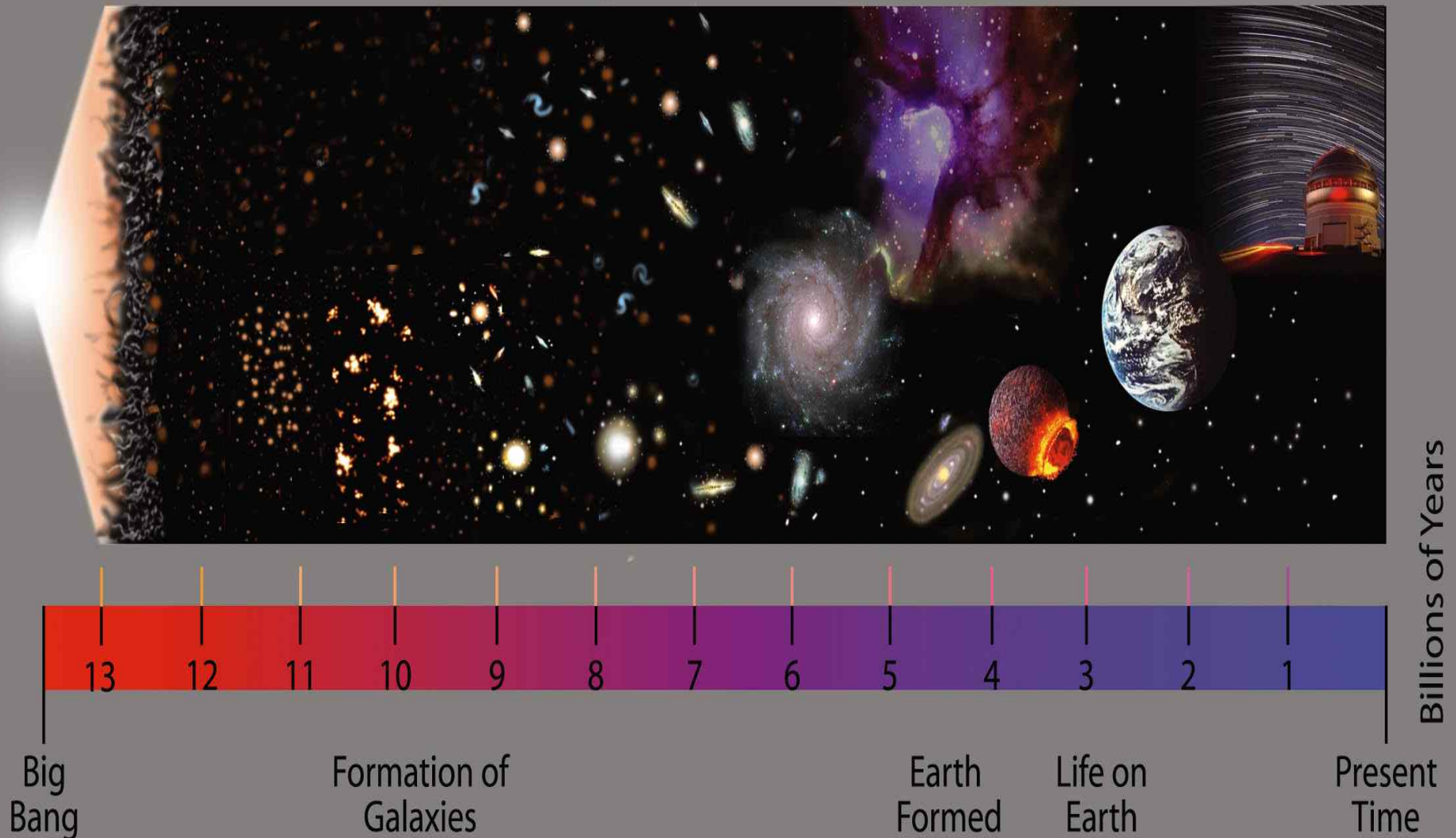
Misturas racêmicas de DNAs e RNAs não conseguem se replicar

Aminoácidos podem servir de catalizadores para produção assimétrica de carboidratos

A homquiralidade é fundamental para que exista a vida!



**Depois de 4.5 Giga-years – O Nosso sistema solar nasceu num universo que tinha 10 Gyears!
Muito da química pré-biótica já poderia estar lá.**

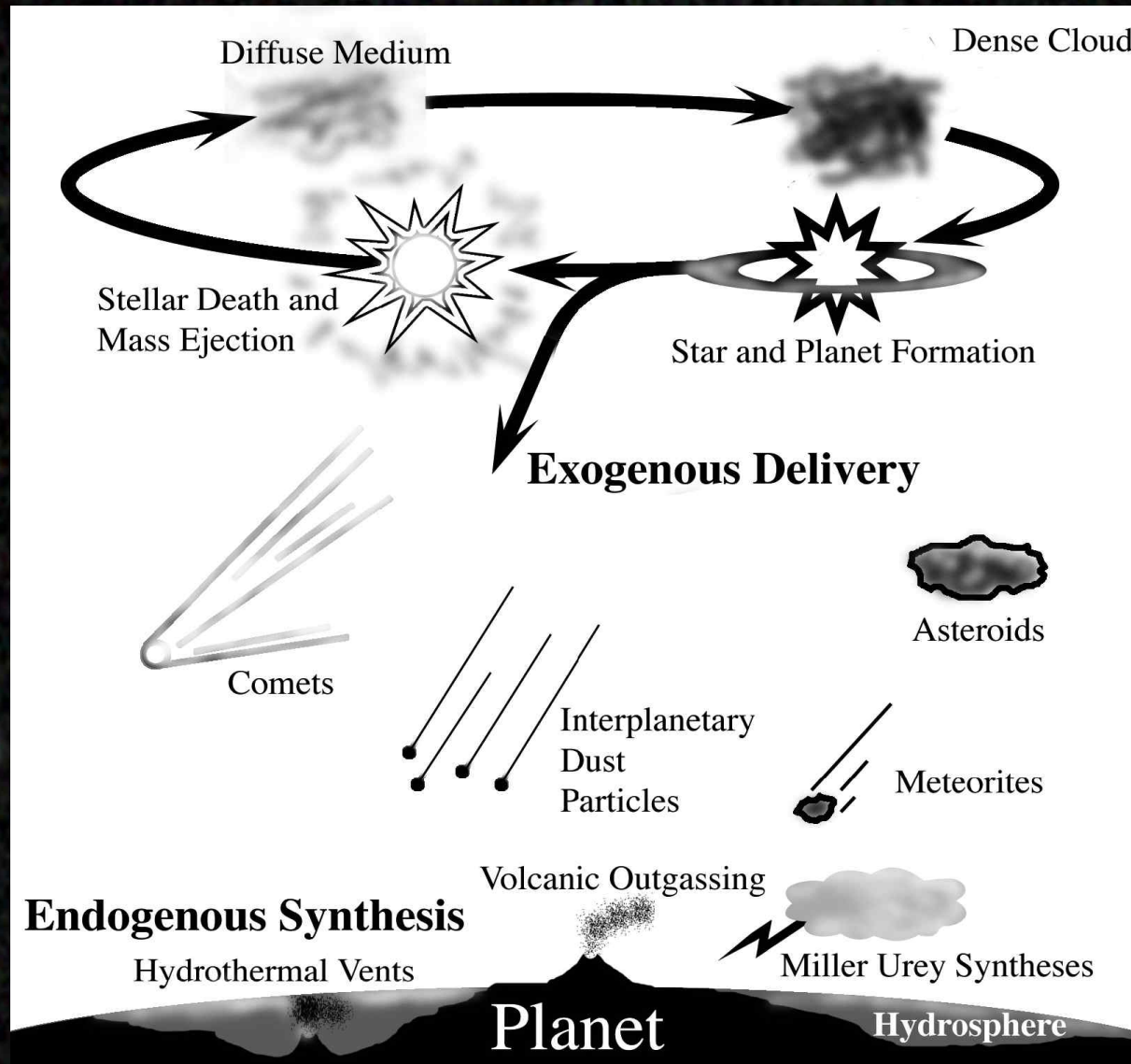


The background of the entire image is the Hubble Ultra Deep Field, a deep-space photograph showing a vast number of galaxies in various shapes and sizes, including spirals, ellipticals, and irregulars, scattered across a black cosmic background. The text is overlaid on this image.

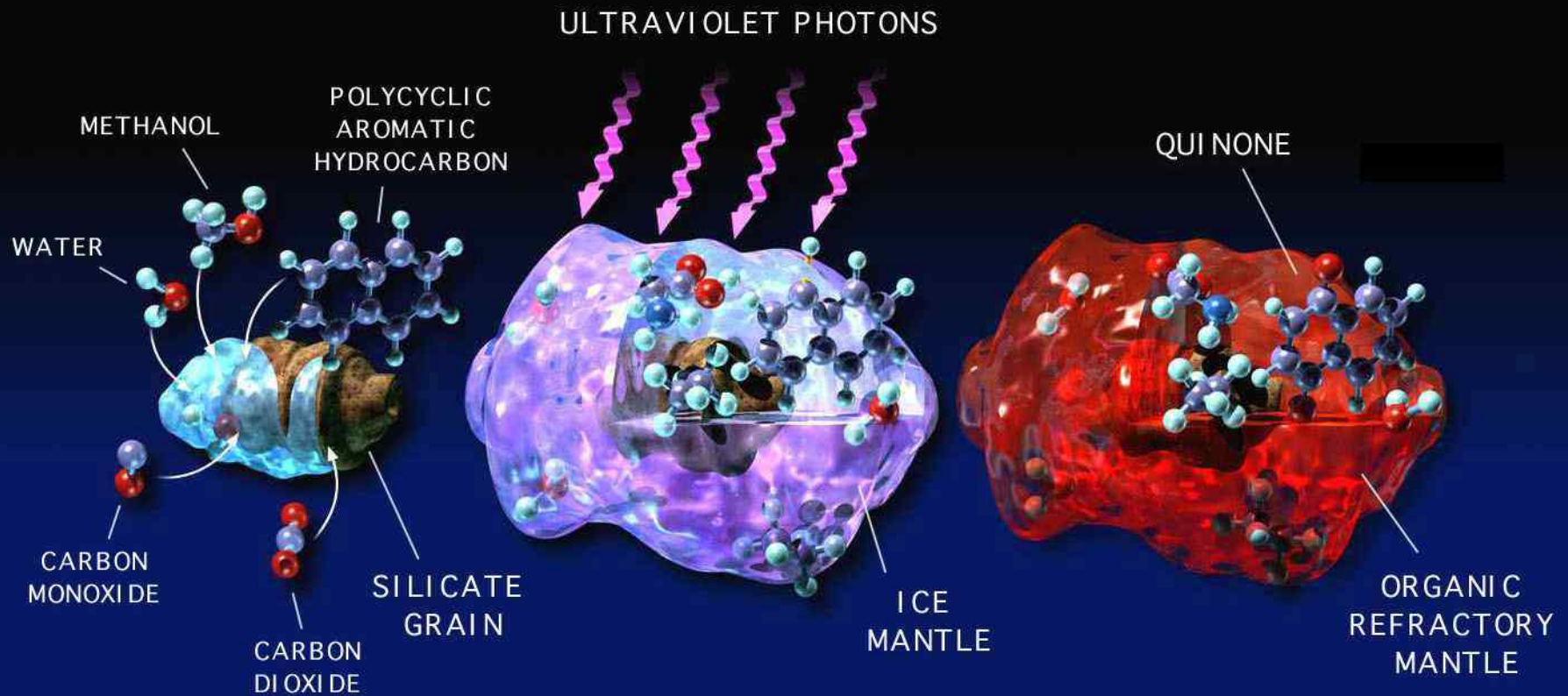
100.000.000.000.000.000.000.000
STARS

Hubble ULTRA deep field

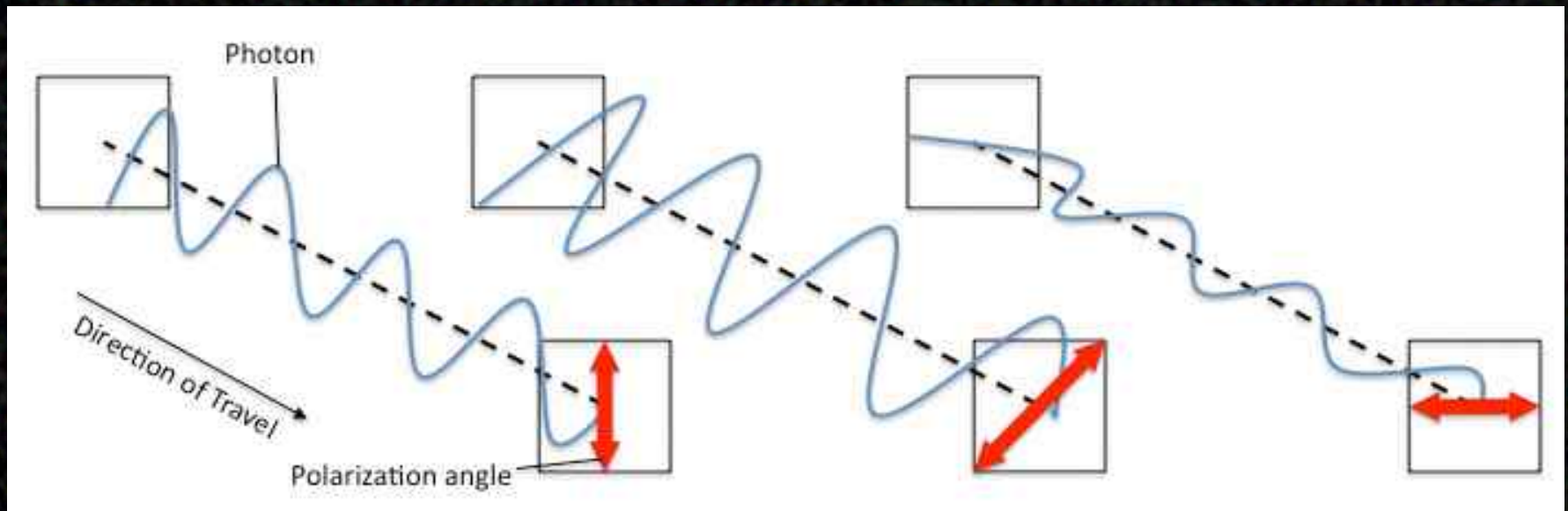
Possíveis caminhos para a química pré-biótica



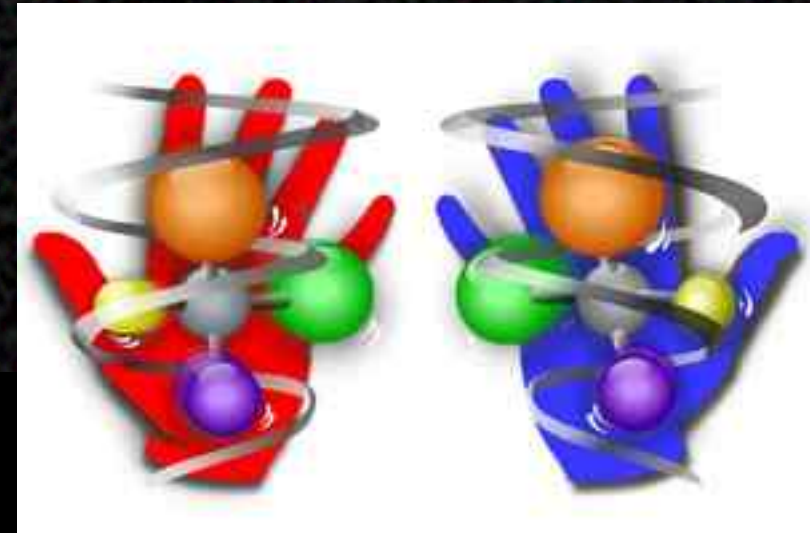
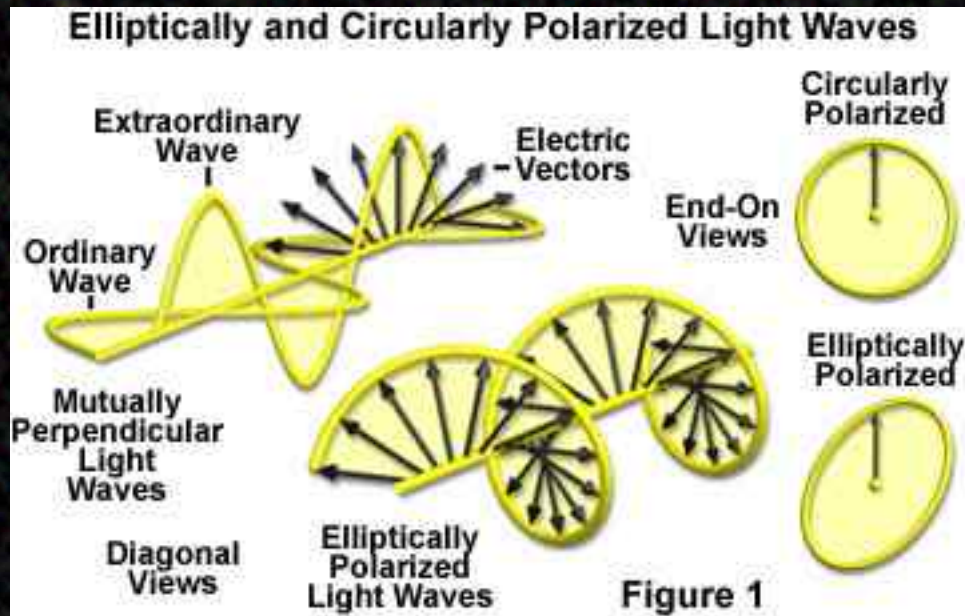
Vamos nos concentrar na produção abiótica no espaço!



Raios-X ou qualquer outro fóton possui uma propriedade importante: POLARIZAÇÃO



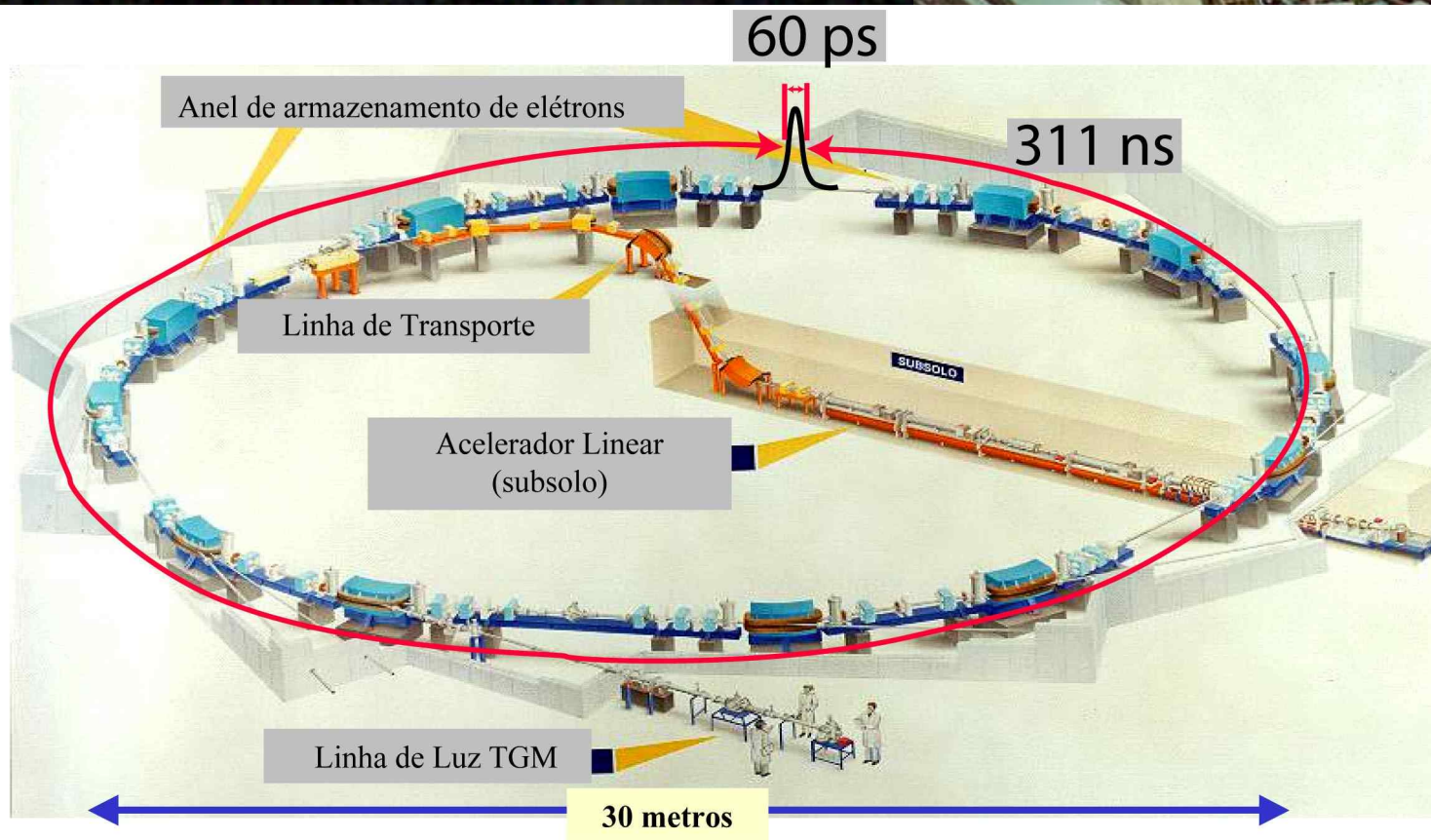
POLARIZAÇÃO Circular de fótons:



**Usaremos como fonte de raios-
x vindo de um acelerador
síncrotron**

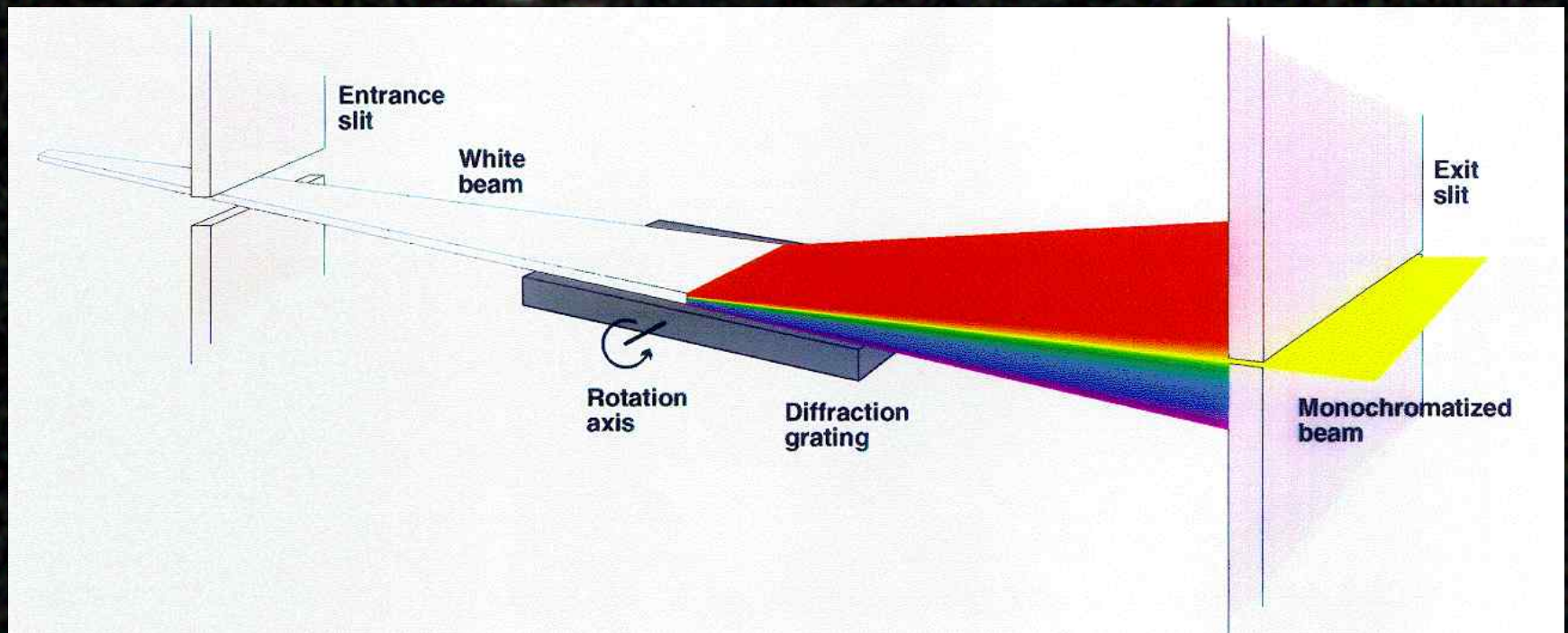
**Fonte muito poderosa (alto
brilho): milhões de anos
acumulados de dose no espaço
são equivalentes a somente
dezenas de horas no síncrotron.**

O LNLS-Campinas



Usualmente temos 147 pacotes separados por 2 ns

Precisamos de um monocromador



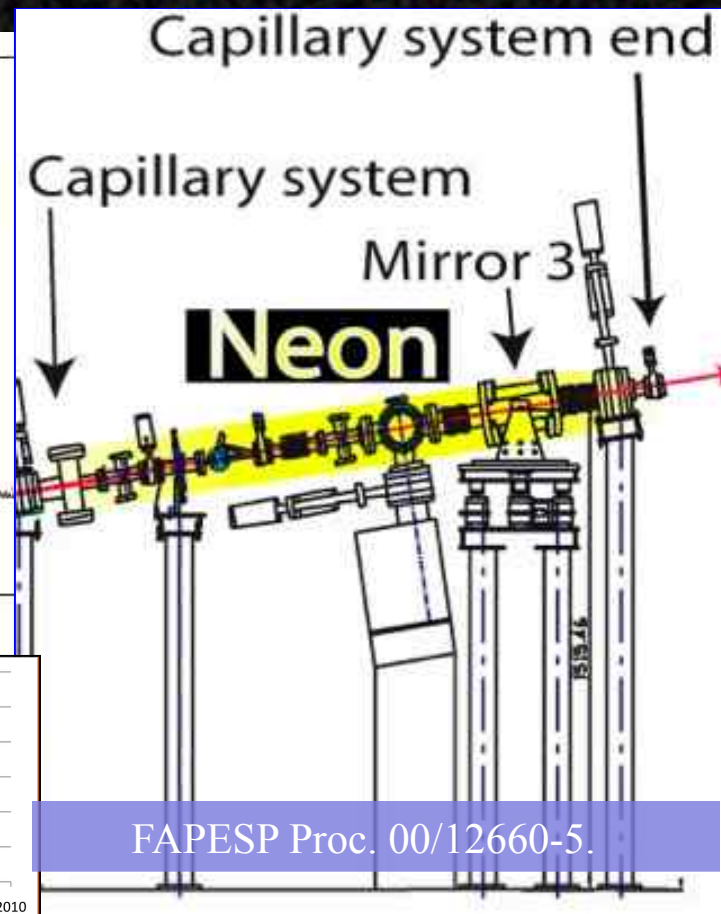
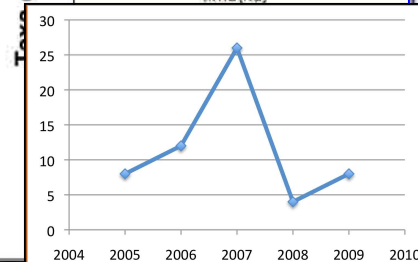
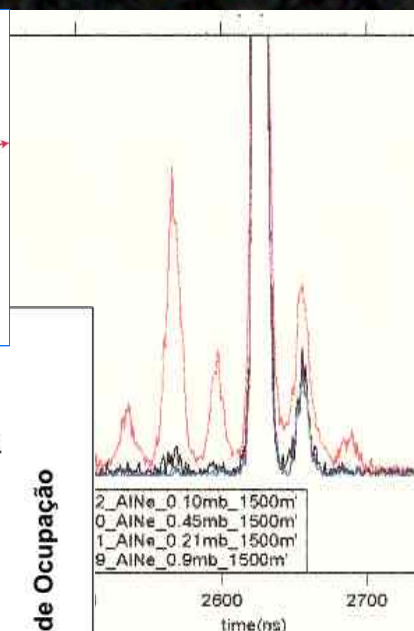
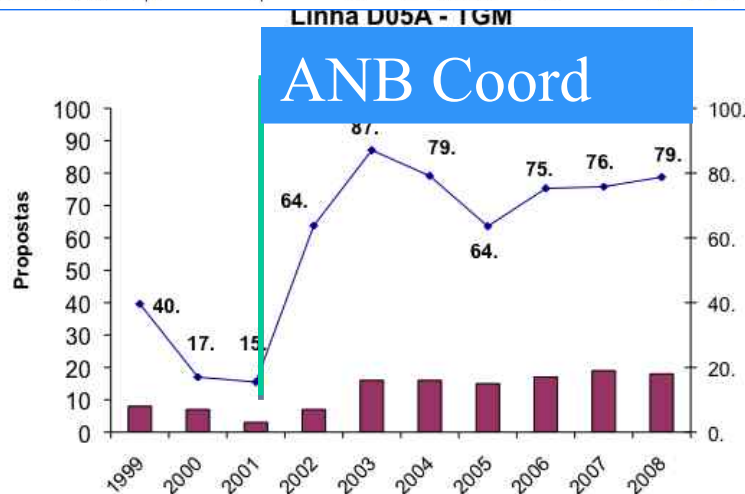
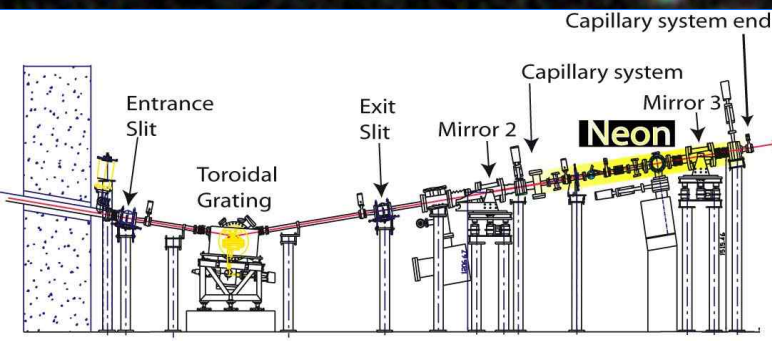
Linha de Luz TGM

Mirhacile

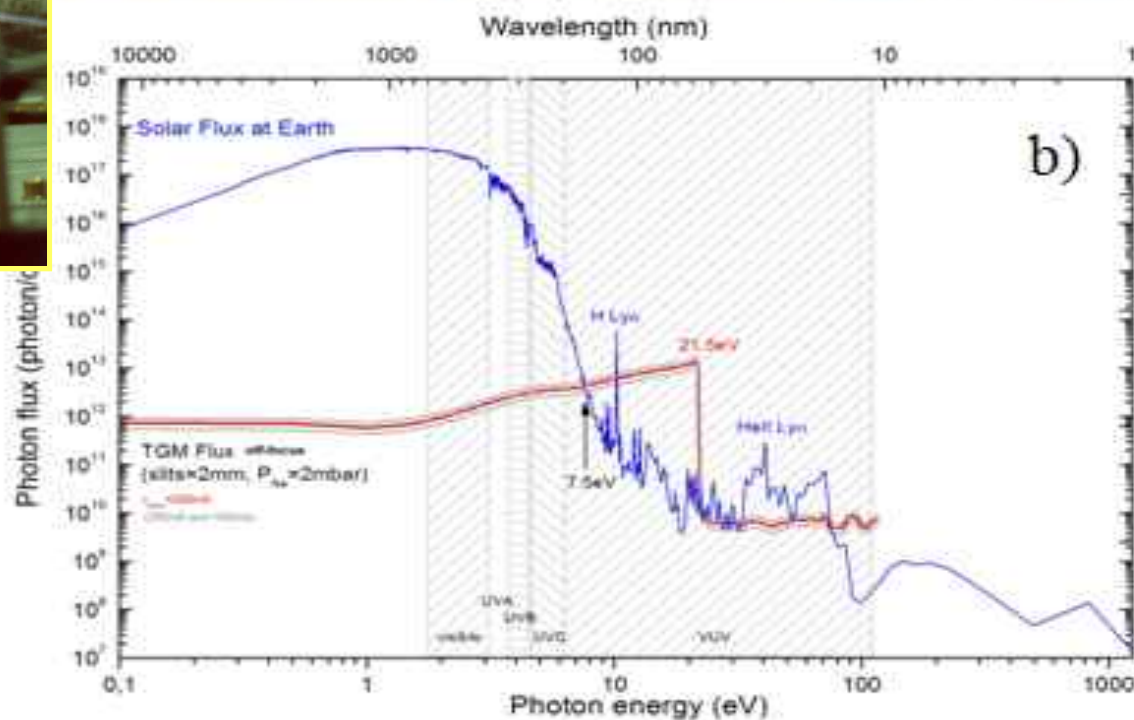
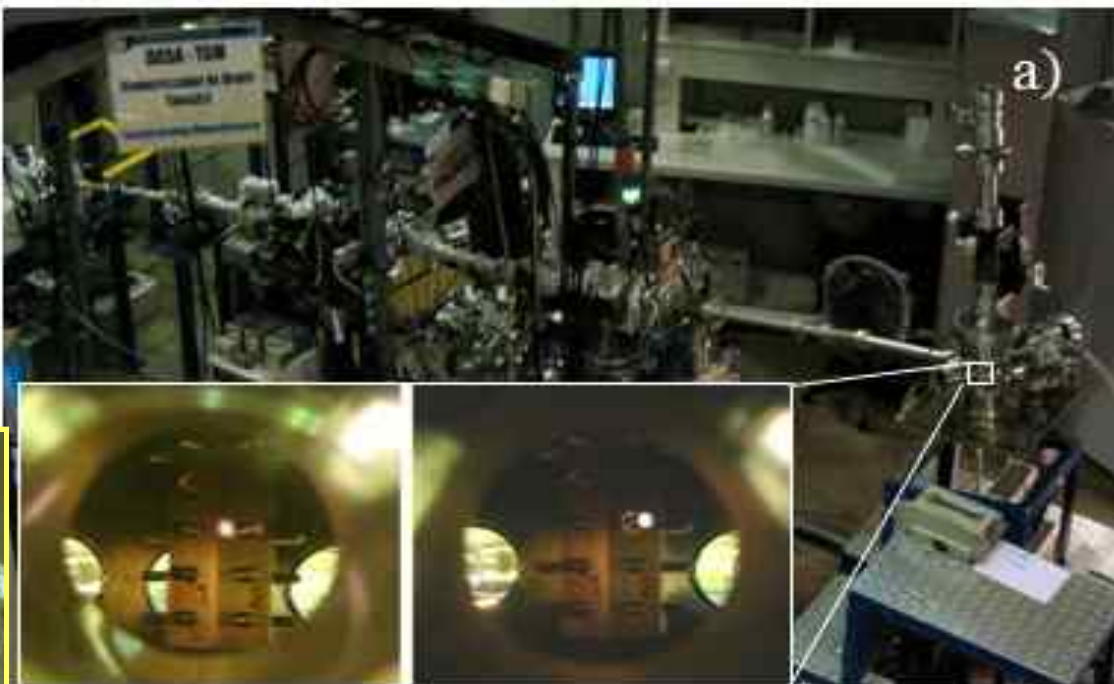
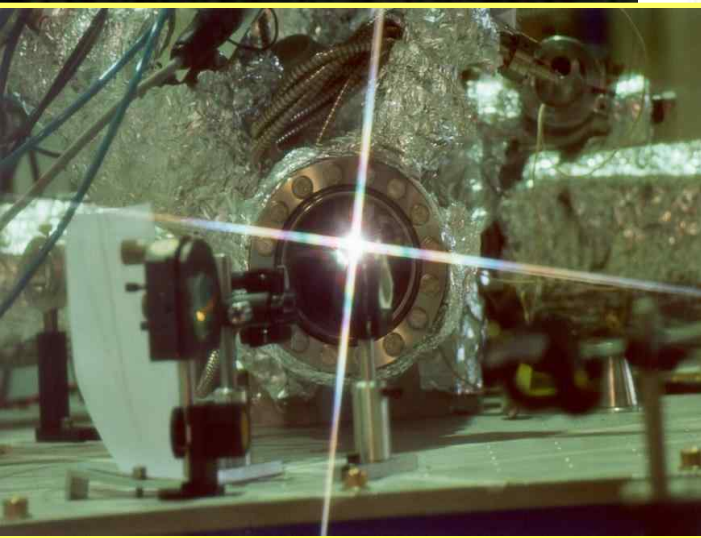
Open Facility (2002- ; 8 years)
TGM: ~40 Projects or 30 group/year
Trained~16 (18) Groups 50 users
trained for TGM al. us. SXS&SGM

MIRror-inserted Harmonic Attenuator
 Calibrator-device with a Long LEgth

Cavasso-Fillho. RL...de Brito AN, Rev
 Scien. Instrum. **78**,115104, (2007)

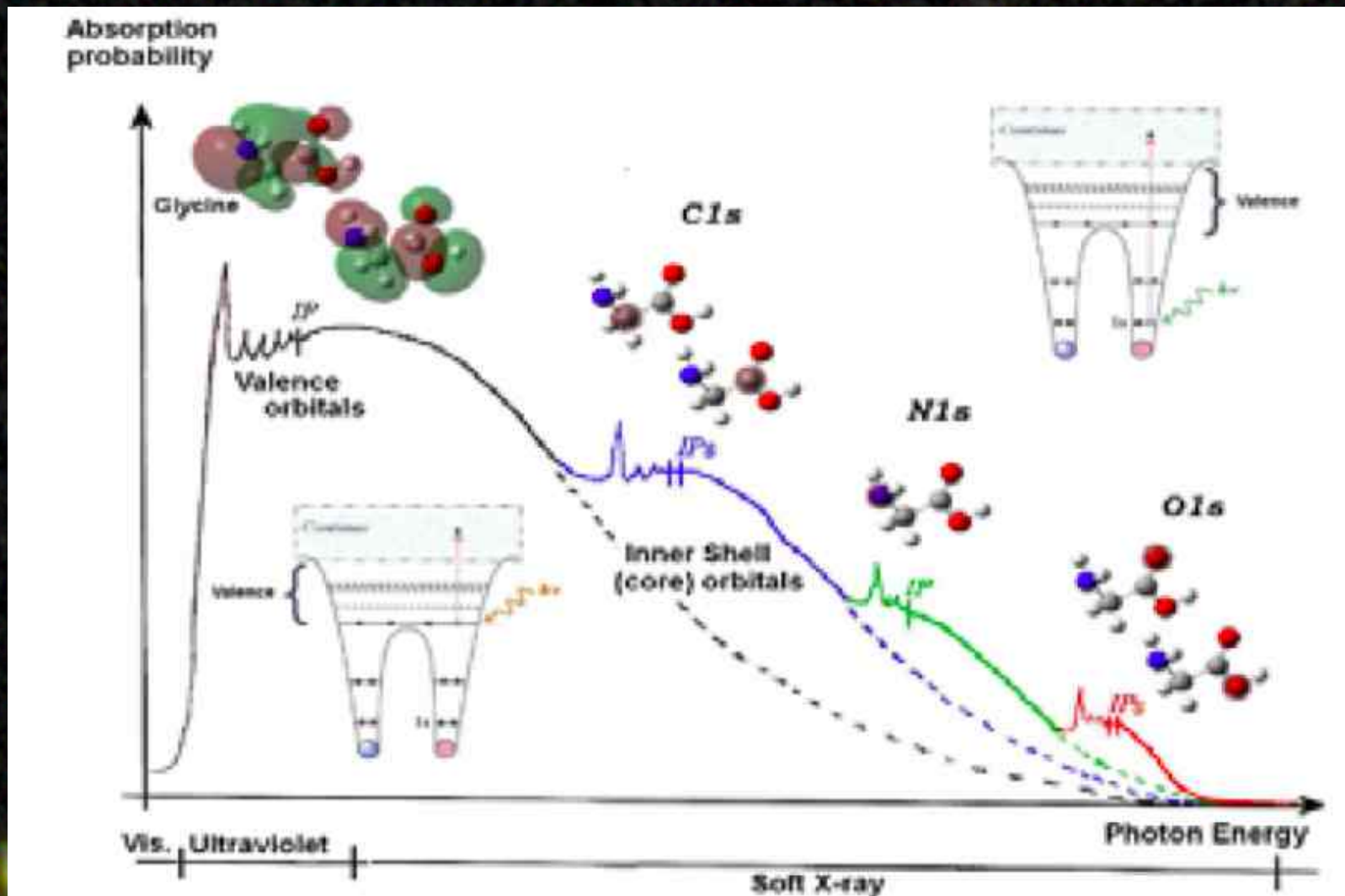


FAPESP Proc. 00/12660-5.



**VUV SOLAR
RADIATION**

VUV não basta raio-x é muito importante



Linha PGM

• Valor total do Projeto incluindo câmaras Experimentais 9MR\$

• Pronex Temático 03-09 como Pesquisador Principal FAPESP
316KU\$ 341KR\$

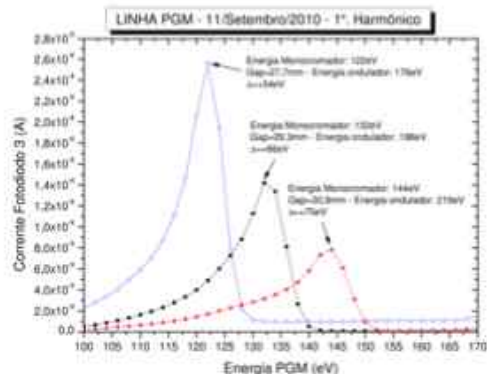
Monitor de posição do feixe II

Câmera de espelhos

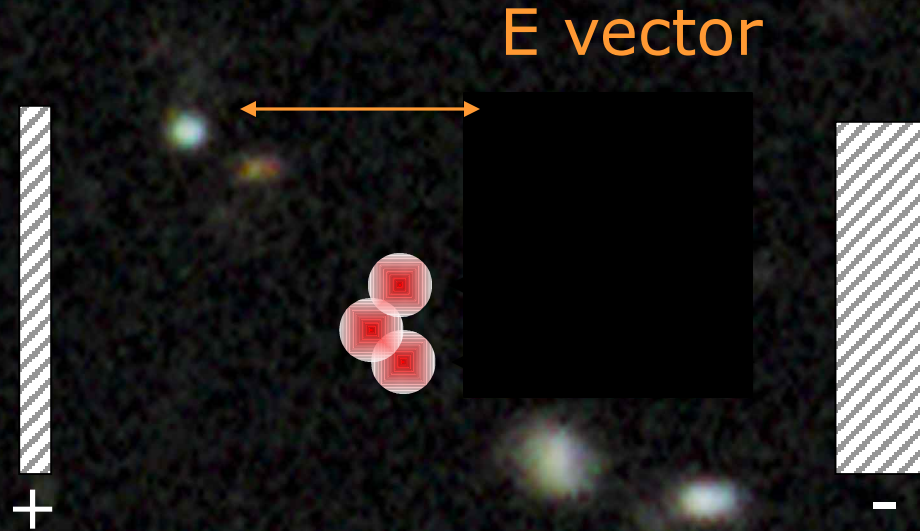
2010

35 m

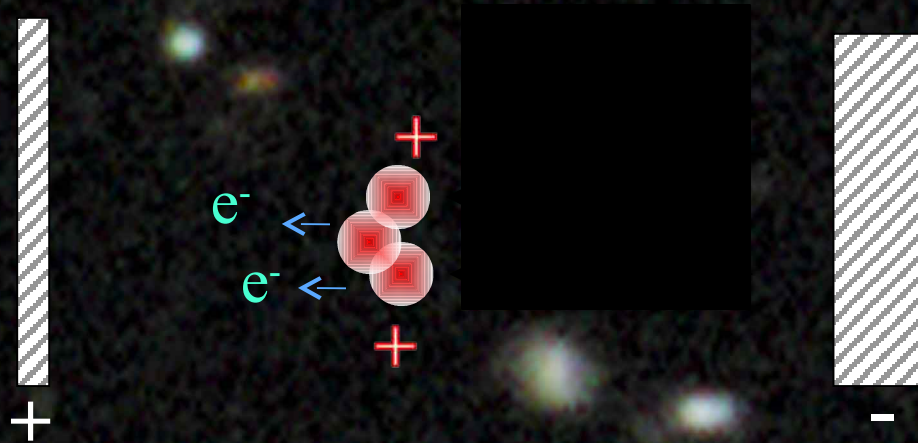
Componentes para bombeamento diferencial



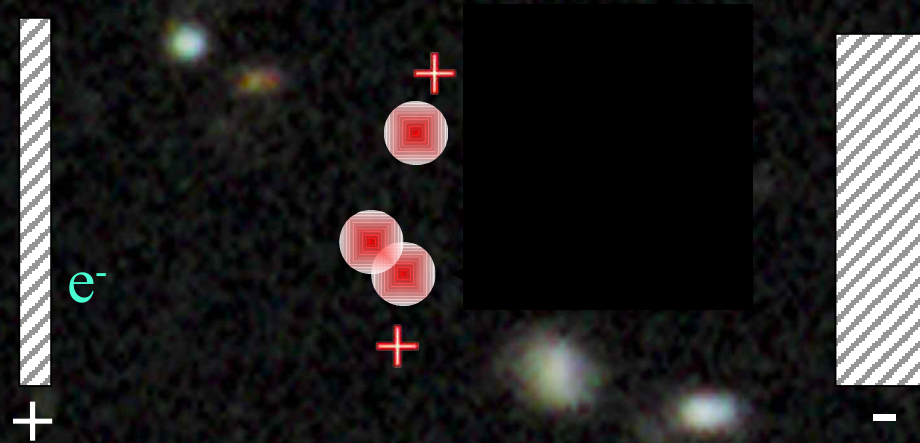
Espectrômetro por Tempo de Voo



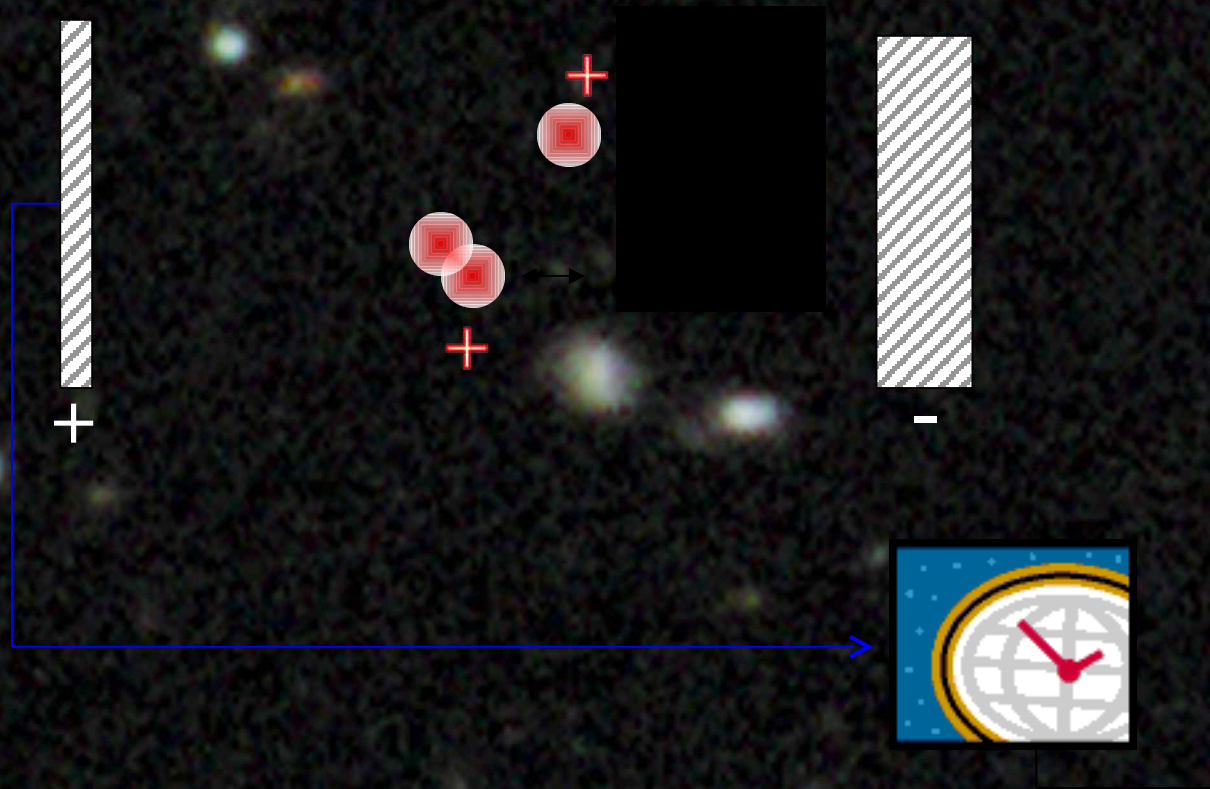
Espectrômetro por Tempo de Voo



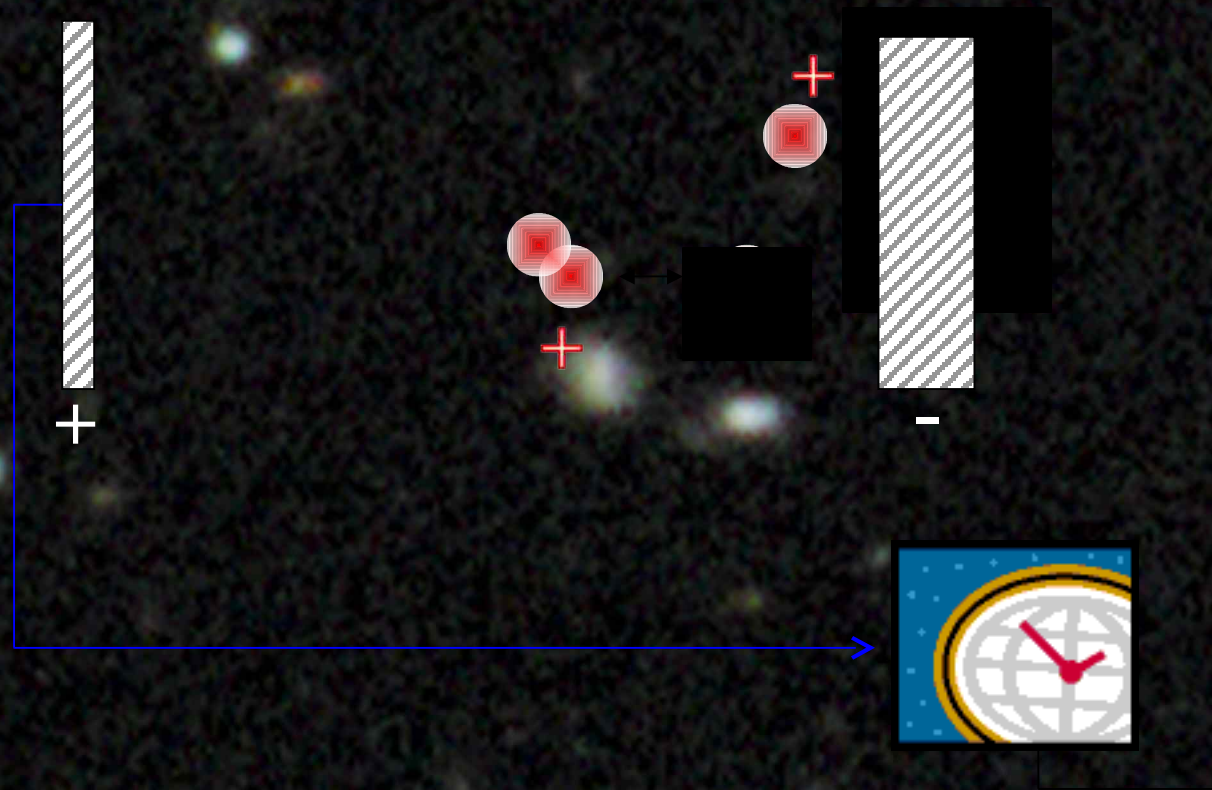
Espectrômetro por Tempo de Voo



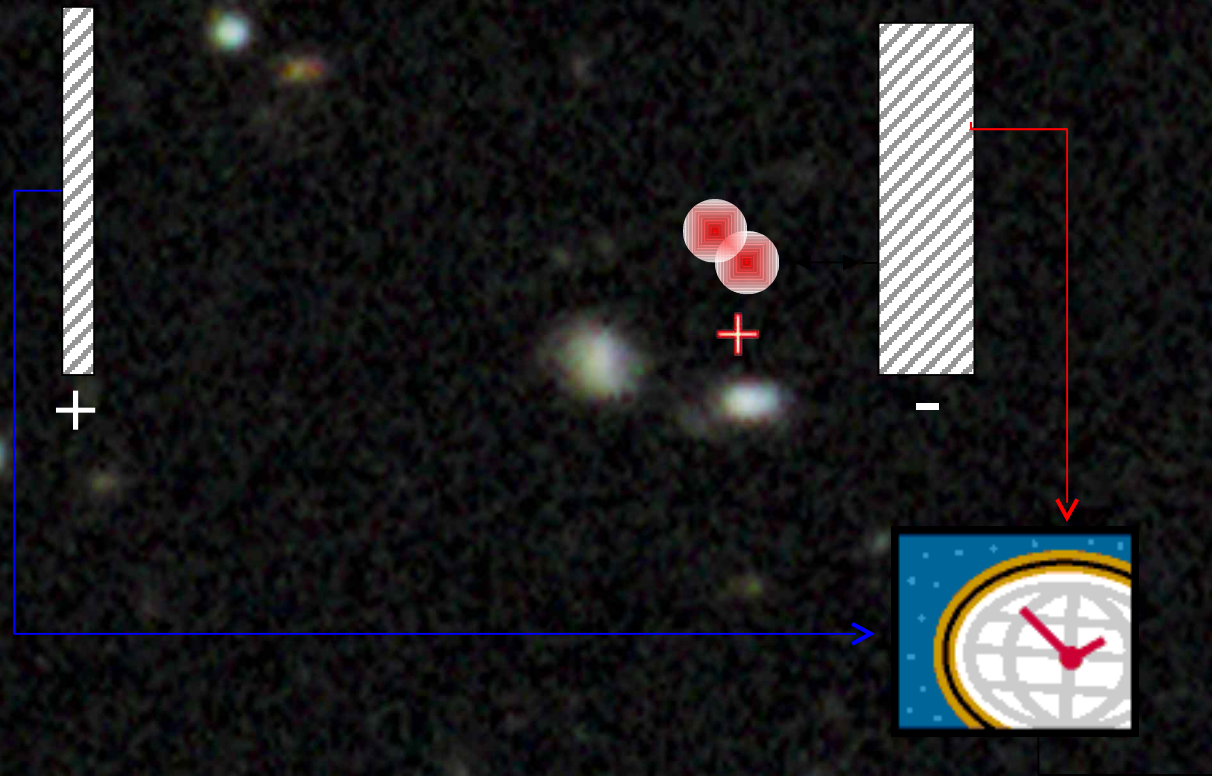
Espectrômetro por Tempo de Voo



Espectrômetro por Tempo de Voo



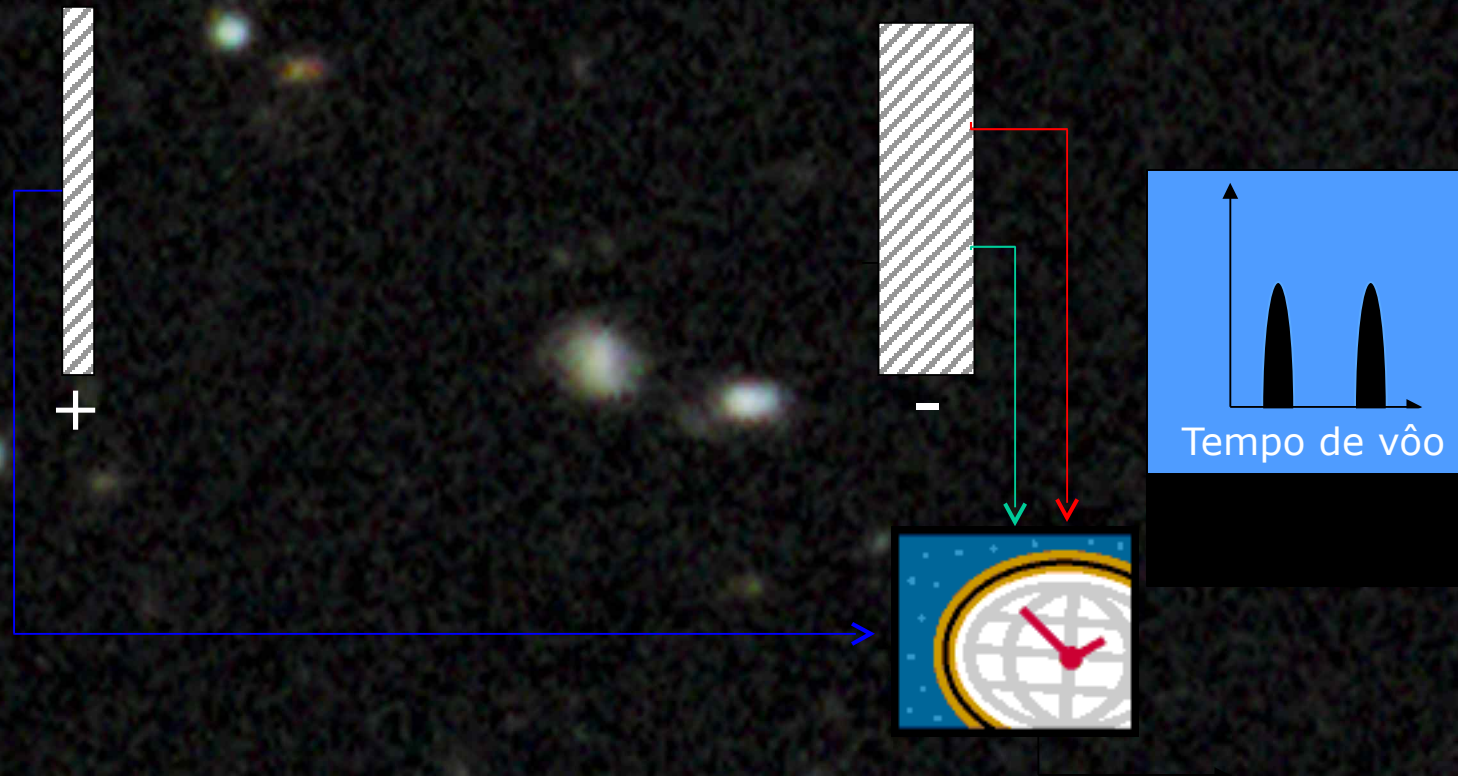
Espectrômetro por Tempo de Voo

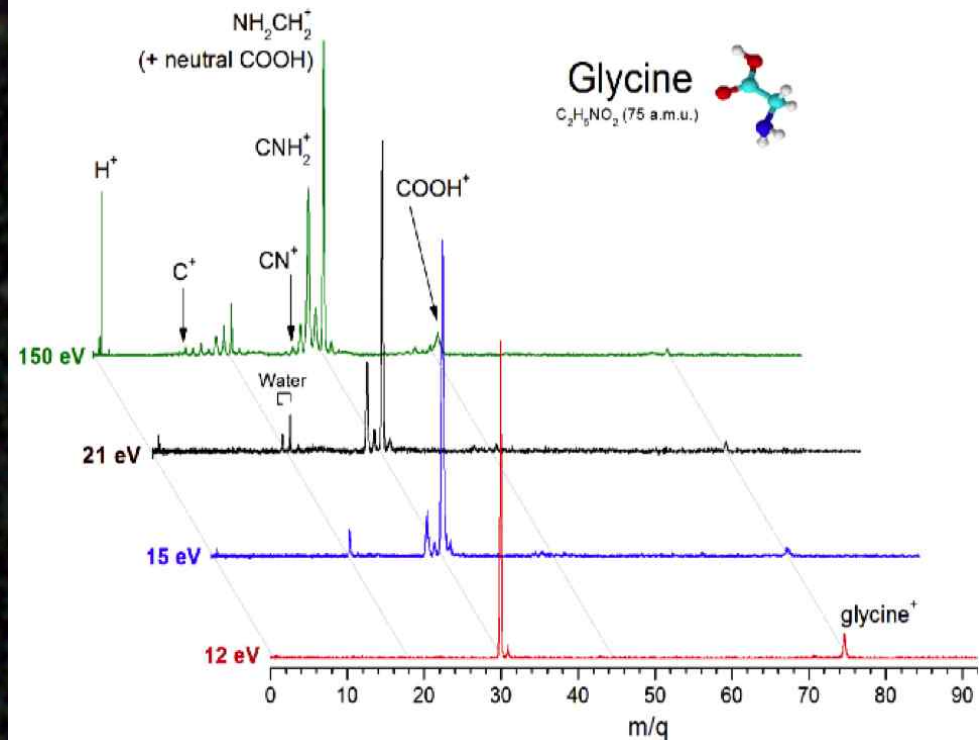
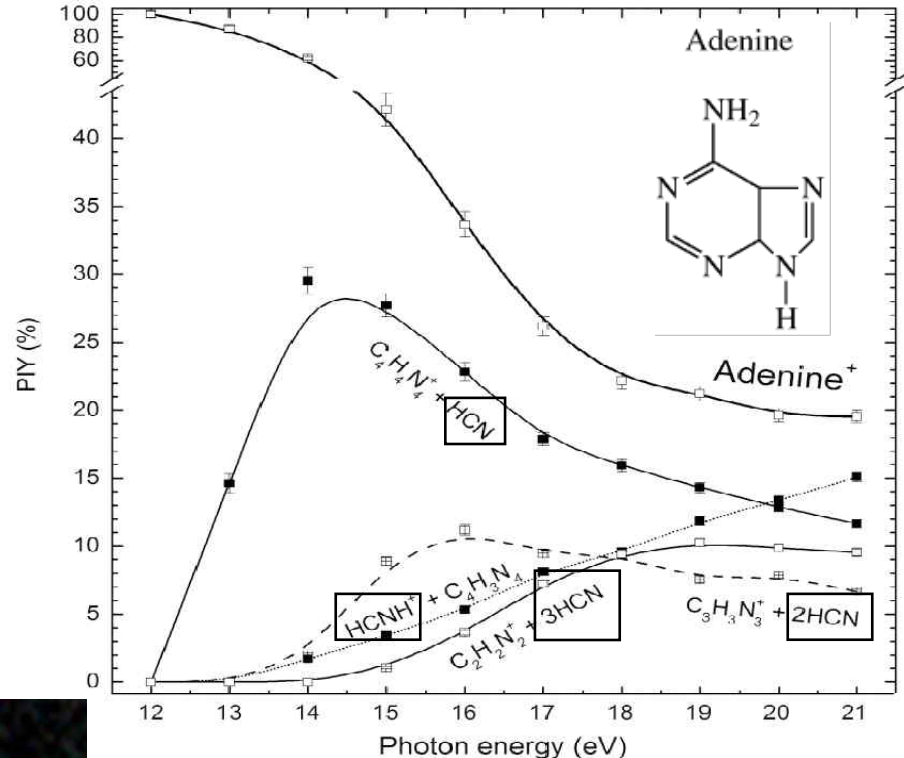


Espectrômetro por Tempo de Voo



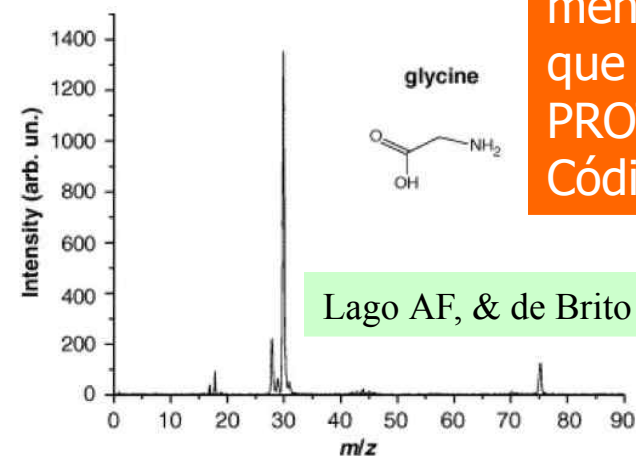
Espectrômetro por Tempo de Vôo





Serg. Pilling & de Brito AN
 Rapid Comm. Mass Spec. **21**,
 3646-3652 (2007)

-Aminoácidos
 fotodegradam em
 energias muito
 menores do que do
 que bases de DNA!!
 PROTEÇÃO extra ao
 Código Genético?



Lago AF, & de Brito AN Chem Phys. **307**, 9 (2004)

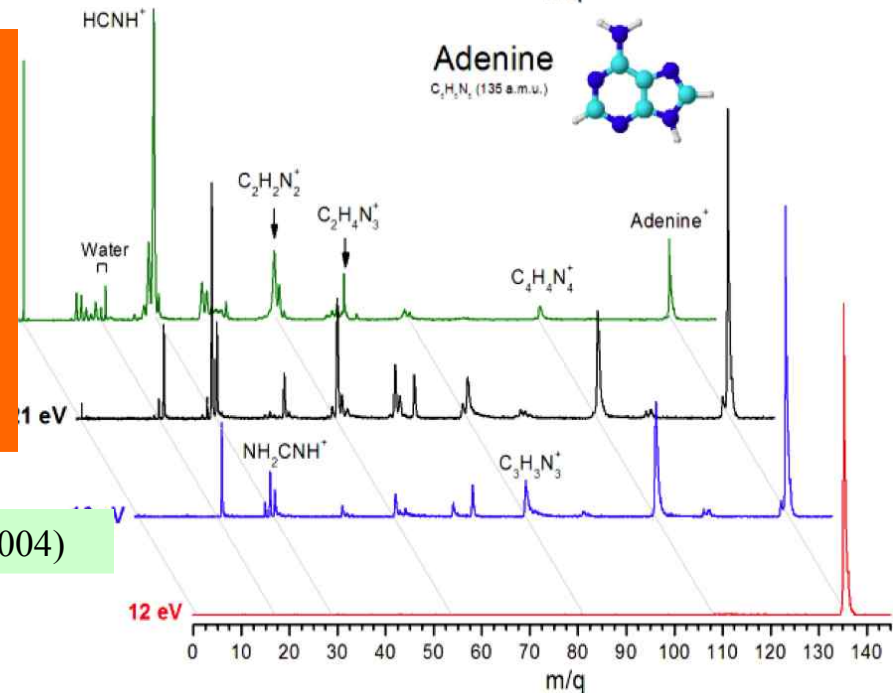
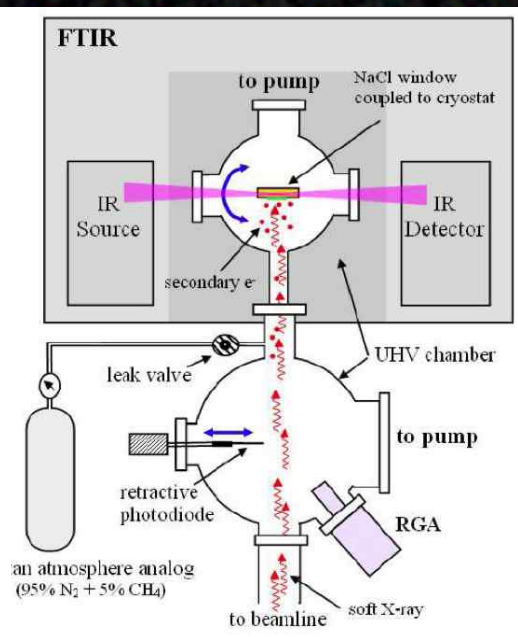
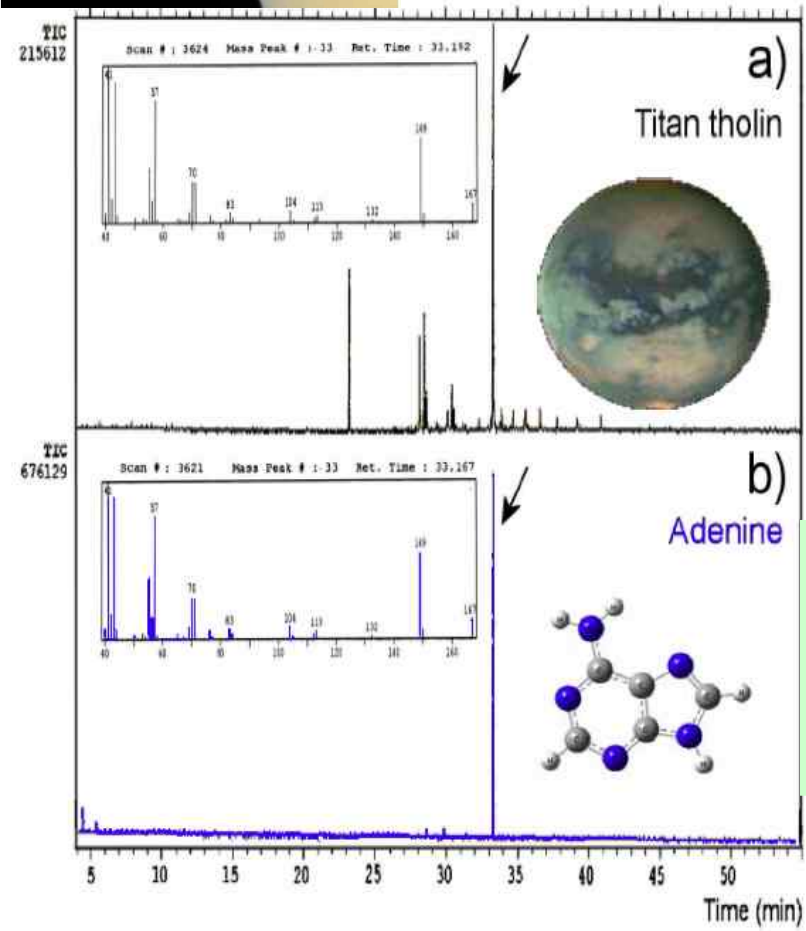
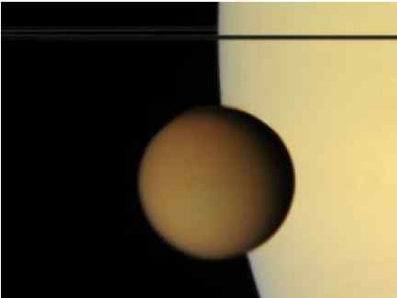


Foto-produção de base de DNA na atmosfera de Titã



Pilling S ...Rittner R
(UNICAMP)&de
Brito AN, JPC A,
113, 11161 (2009),

- 73 hs de irradiação com luz sincrotron!
- ~ 7 milhões de anos de raios X solares em Titã
- Adenina: formação promovida por elétrons secundários produzido a partir da dissociação de moléculas por fótons de raios X.

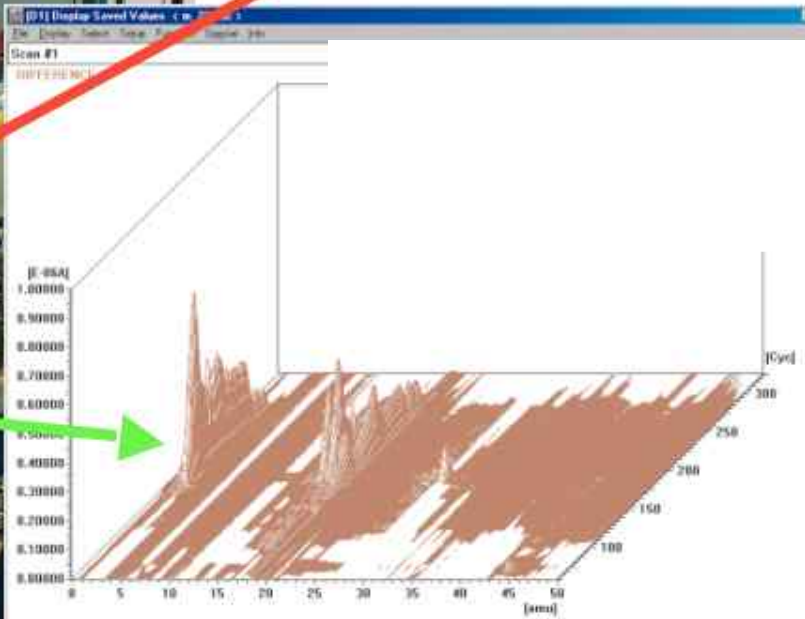
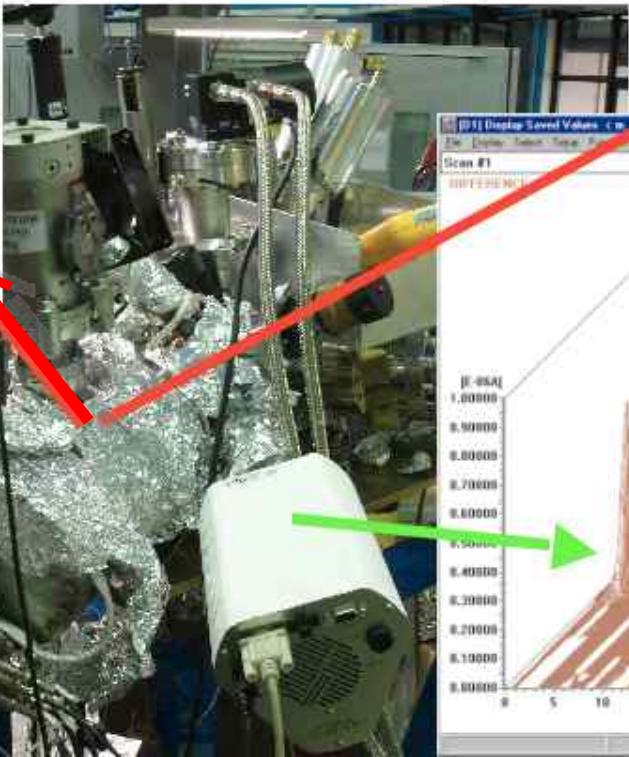
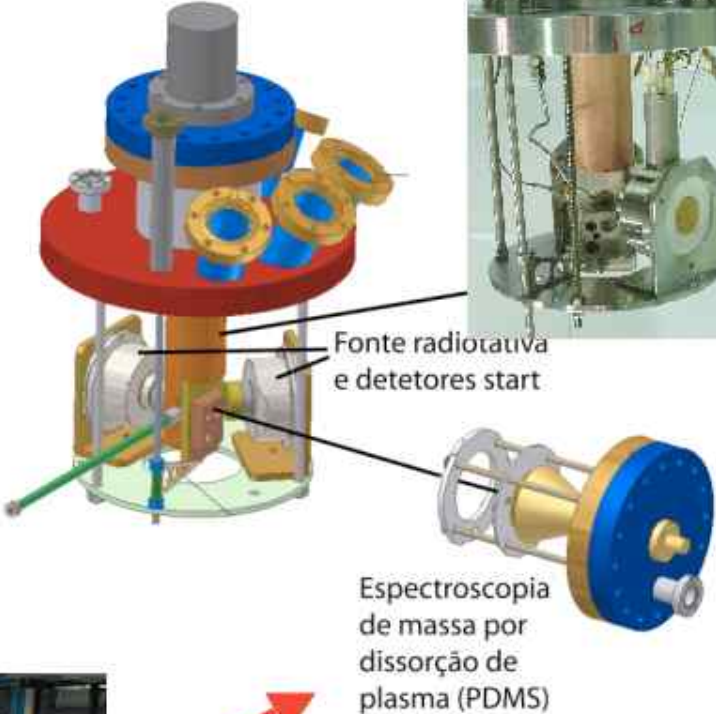
- Produção de base de DNA em simulação de lua de Saturno:

Comet “simulator” at LNLS

Sample	ions	LS
H ₂ O	✓	✓
	✓	✓
CO ₂		
H ₂ O +	✓	
CO ₂		
CO	✓	✓
NH ₃	✓	
NH ₃ + CO	✓	✓
NH ₃ + CO + H ₂ O	✓	✓

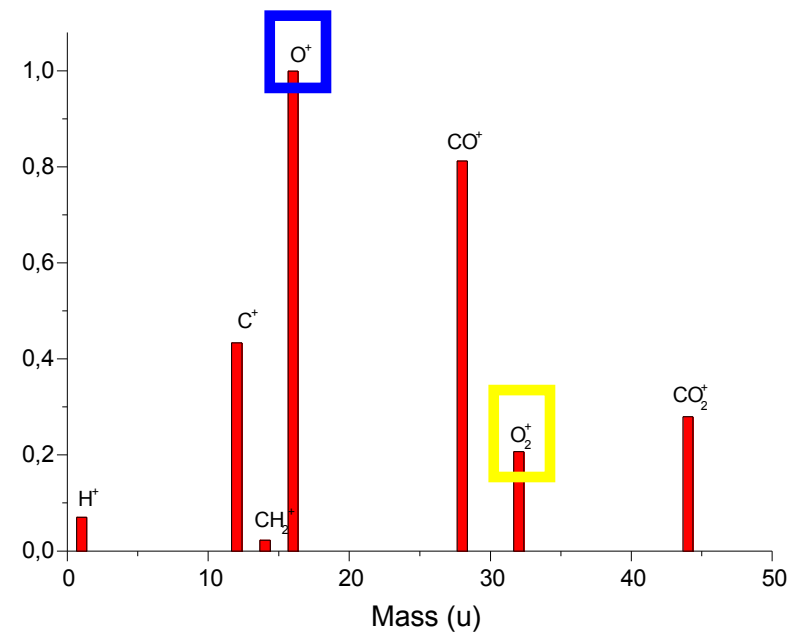
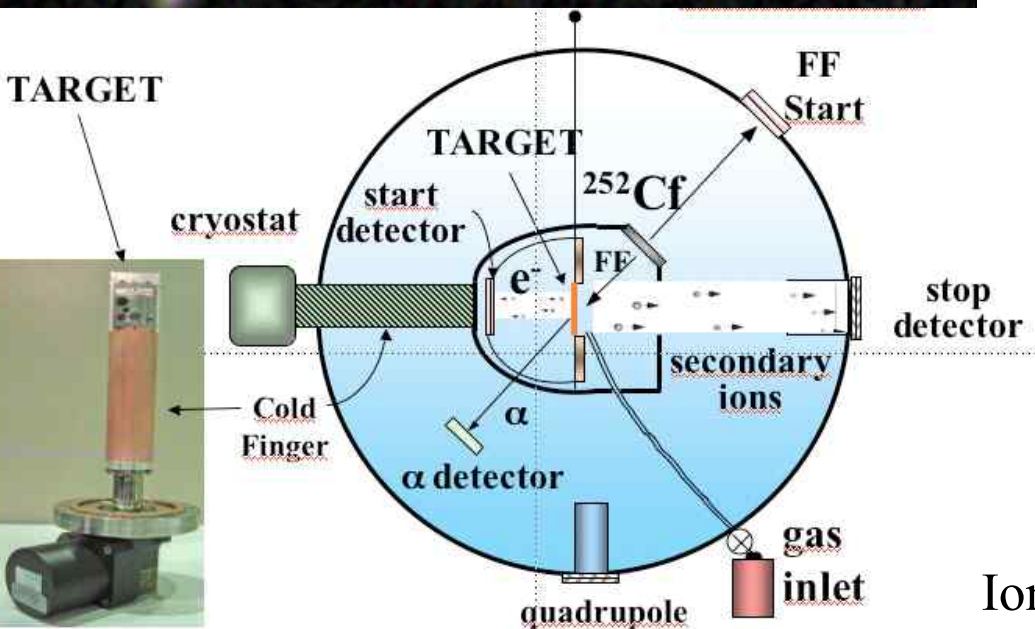
In-situ probing of the formation of amino acid precursors

molecules



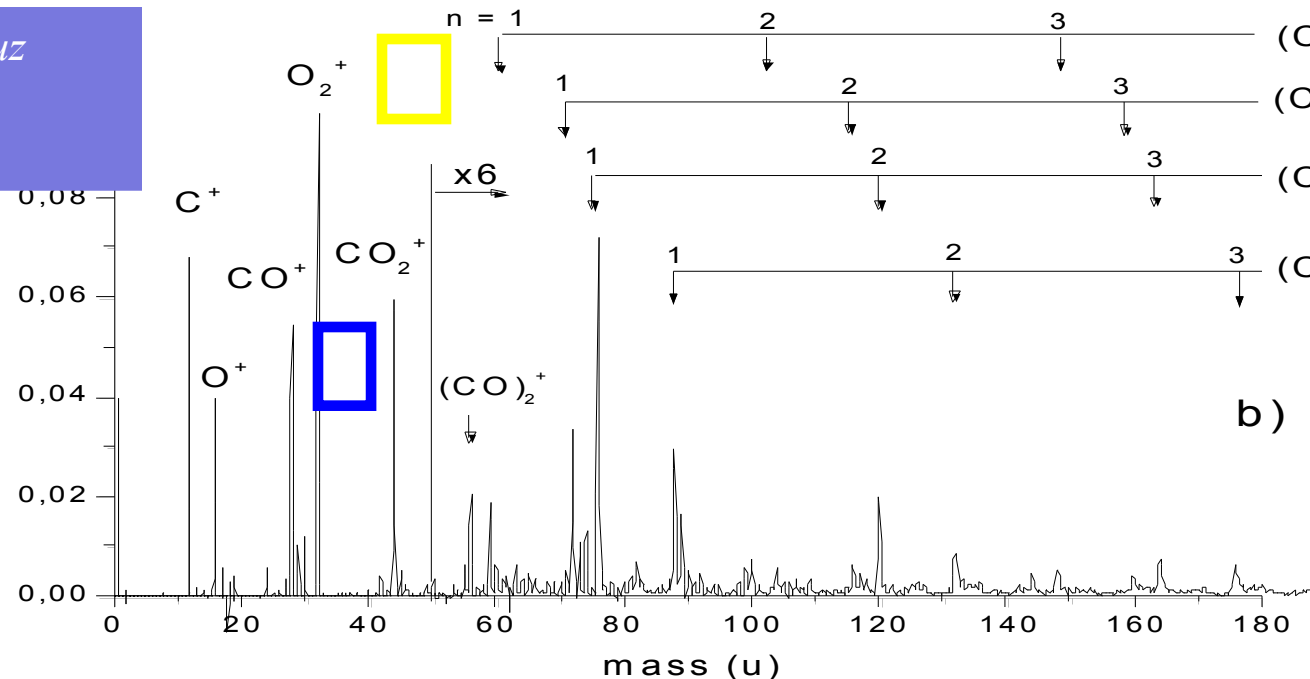
ESA, Alfred Vokshodt (Institut d'Astrophysique de Paris, CNRS, France), NASA

PDMS Chamber Oil



Ions impact O₂ LARGEST INTENSITY

- Projeto "Origem da vida e Luz síncrotron: Simulação....."
- FAPESP Proc. 00/12660-5.)



Panspermia hypothesis



Ejection



Transport in space

Density: $1 - 10^6$
molecules.cm⁻³

Pressure 10^{-17} atm

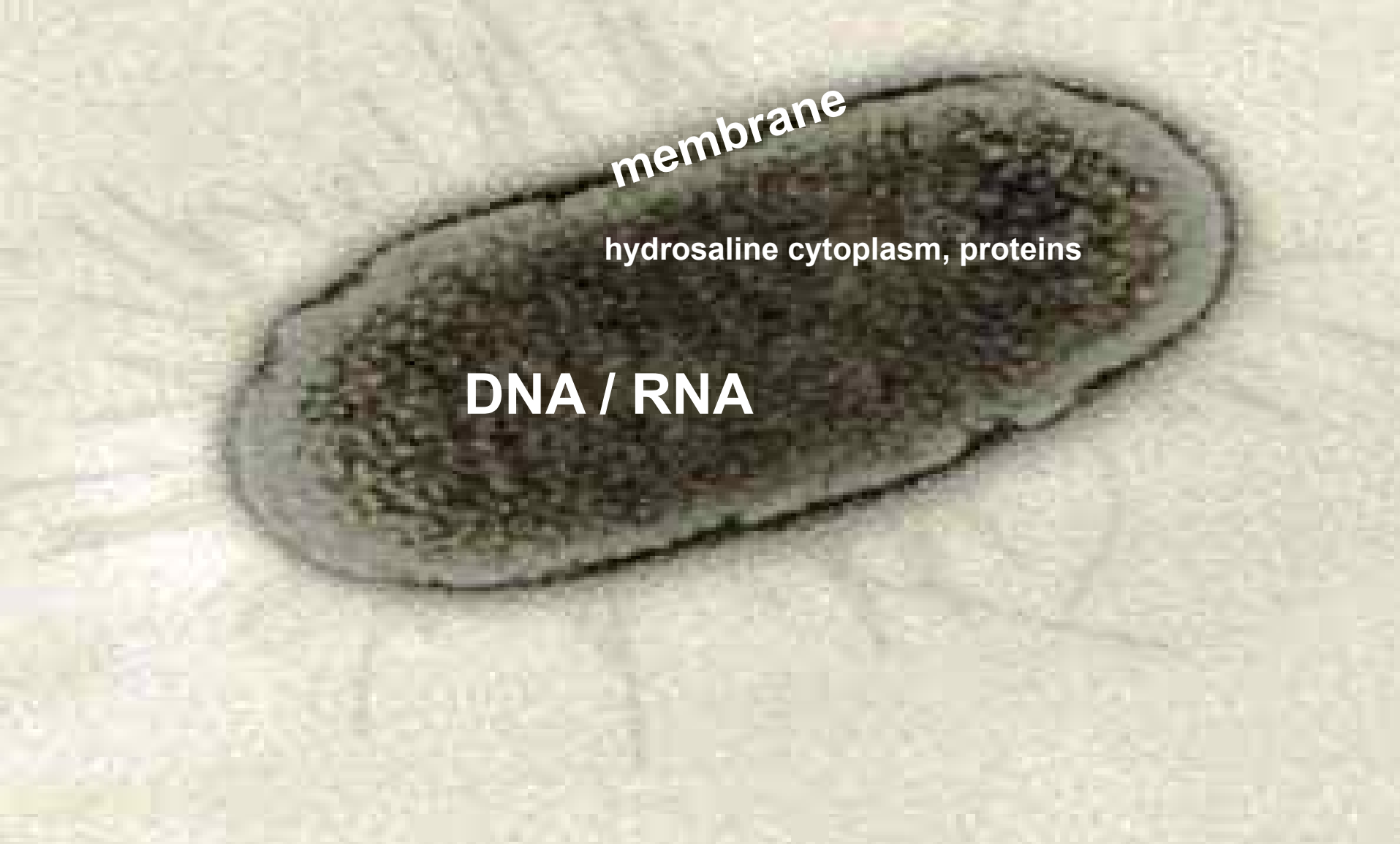
Temperature = $0 - 10^2$ K

UV Radiation:
 $122.3 \text{ J.m}^{-2}.\text{s}^{-1}$



Reentry

Basic structure of a bacterial cell



Apollo Programme (*Streptomyces*, *Bacillus subtilis*)



Table 1 Classification and examples of extremophiles

Environmental parameter	Type	Definition	Examples
Temperature	Hyperthermophile	Growth >80 °C	<i>Pyrolobus fumarii</i> , 113 °C
	Thermophile	Growth 60–80 °C	<i>Synechococcus lividis</i>
	Mesophile	15–60 °C	<i>Homo sapiens</i>
	Psychrophile	<15 °C	<i>Psychrobacter</i> , some insects
Radiation			<i>Deinococcus radiodurans</i>
Pressure	Barophile	Weight-loving	Unknown
	Piezophile	Pressure-loving	For microbe, 130 MPa
Gravity	Hypergravity	>1g	None known
	Hypogravity	<1g	None known
Vacuum		Tolerates vacuum (space devoid of matter)	Tardigrades, insects, microbes, seeds
Desiccation	Xerophiles	Anhydrobiotic	<i>Artemia salina</i> ; nematodes, microbes, fungi, lichens
Salinity	Halophile	Salt-loving (2–5 M NaCl)	Halobacteriaceae, <i>Dunaliella salina</i>
pH	Alkaliphile	pH > 9	<i>Natronobacterium</i> , <i>Bacillus firmus</i> OF4, <i>Spirulina</i> spp. (all pH 10.5)
	Acidophile	low pH-loving	<i>Cyanidium caldarium</i> , <i>Ferroplasma</i> sp. (both pH 0)
Oxygen tension	Anaerobe	Cannot tolerate O ₂	<i>Methanococcus jannaschii</i>
	Microaerophile	Tolerates some O ₂	<i>Clostridium</i>
	Aerobe	Requires O ₂	<i>H. sapiens</i>
Chemical extremes	Gases		<i>C. caldarium</i> (pure CO ₂)
	Metals	Can tolerate high concentrations of metal (metalotolerant)	<i>Ferroplasma acidarmanus</i> (Cu, As, Cd, Zn); <i>Ralstonia</i> sp. CH34 (Zn, Co, Cd, Hg, Pb)



Deinococcus radiodurans:

AN **EXTREME** EXTREMOPHILE

NON-SPORE forming species (always viable)

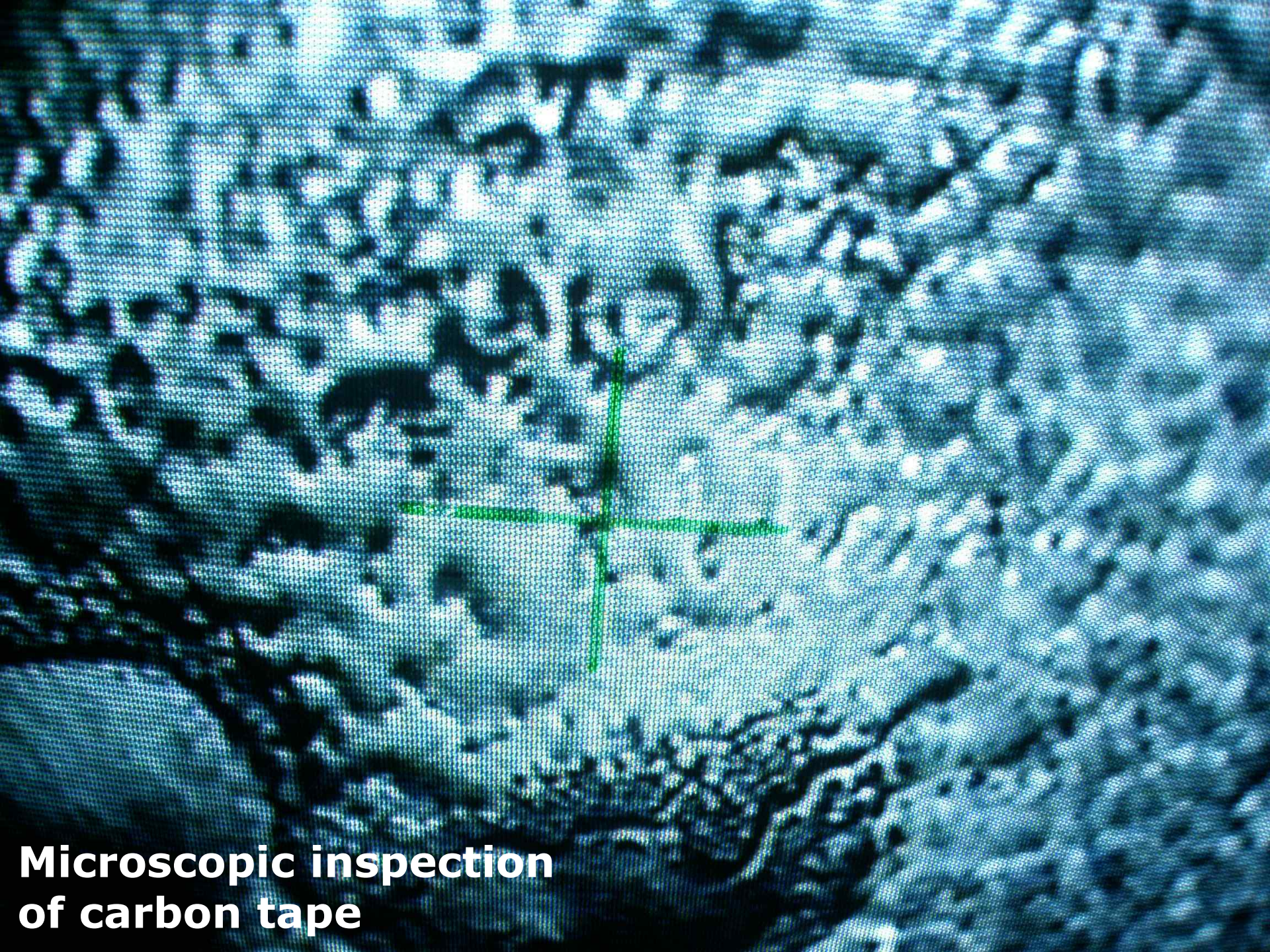
High content of anti-oxidant compounds (carotene and Mn^{++})

Impressive mechanisms to repair DNA damage !!

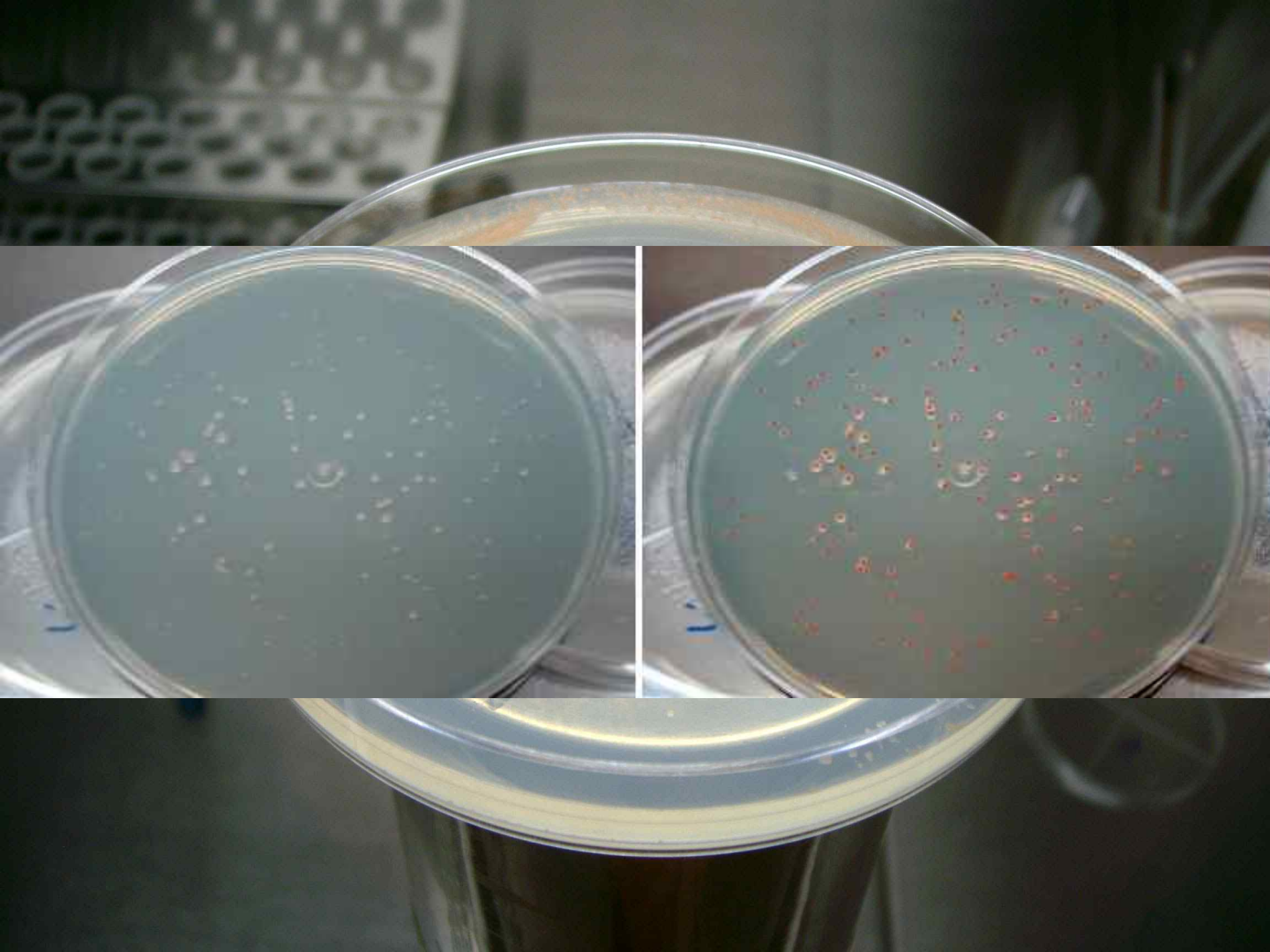
still under division after 15,000 Gy 😊

Humans = 4 Gy !!!!! ☠️ **1 Gray = 1 J/kg.**

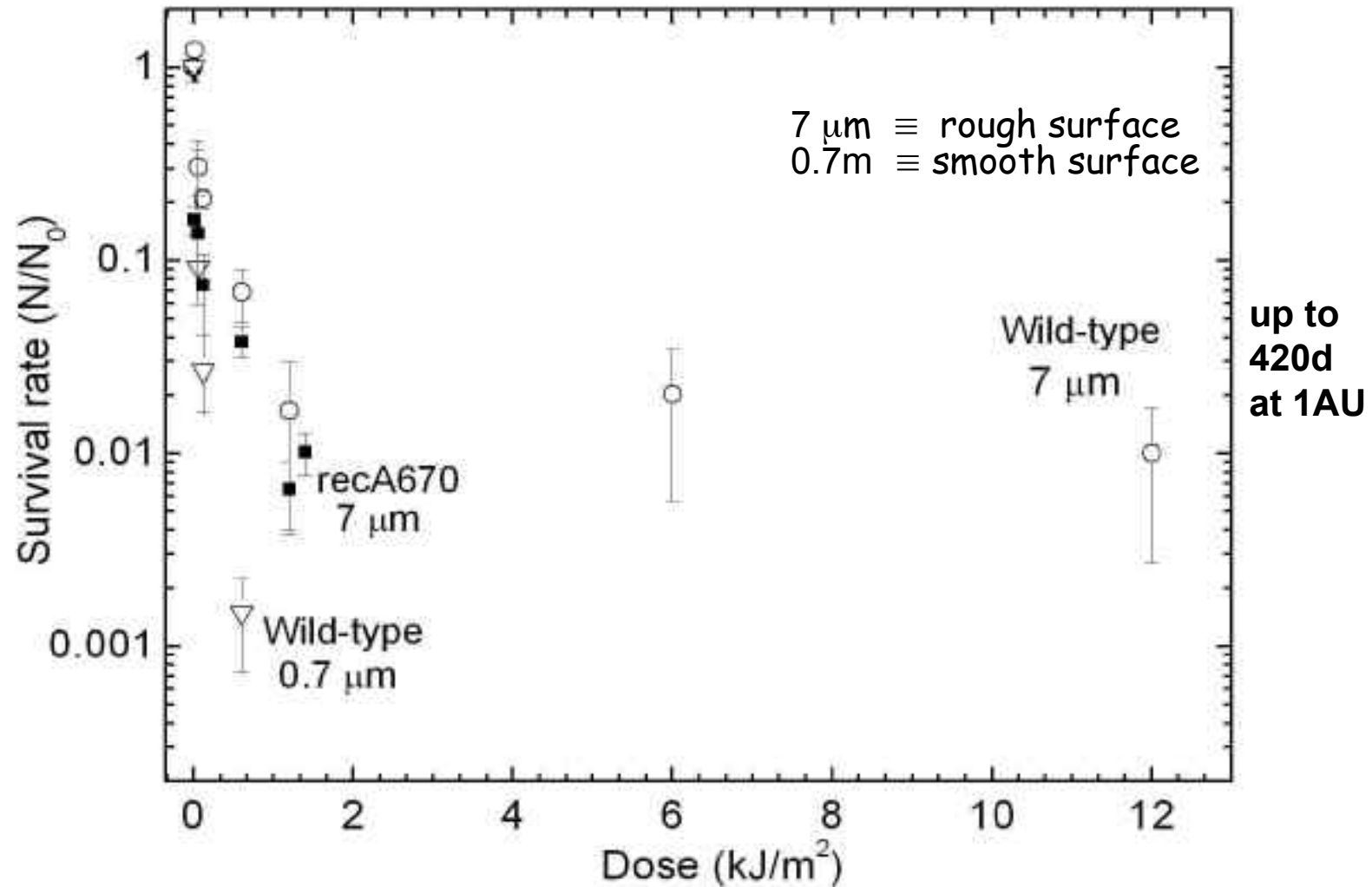
- “minimal” biological organization inoculated among different planets ??
- (if it ever could survive outer space conditions)

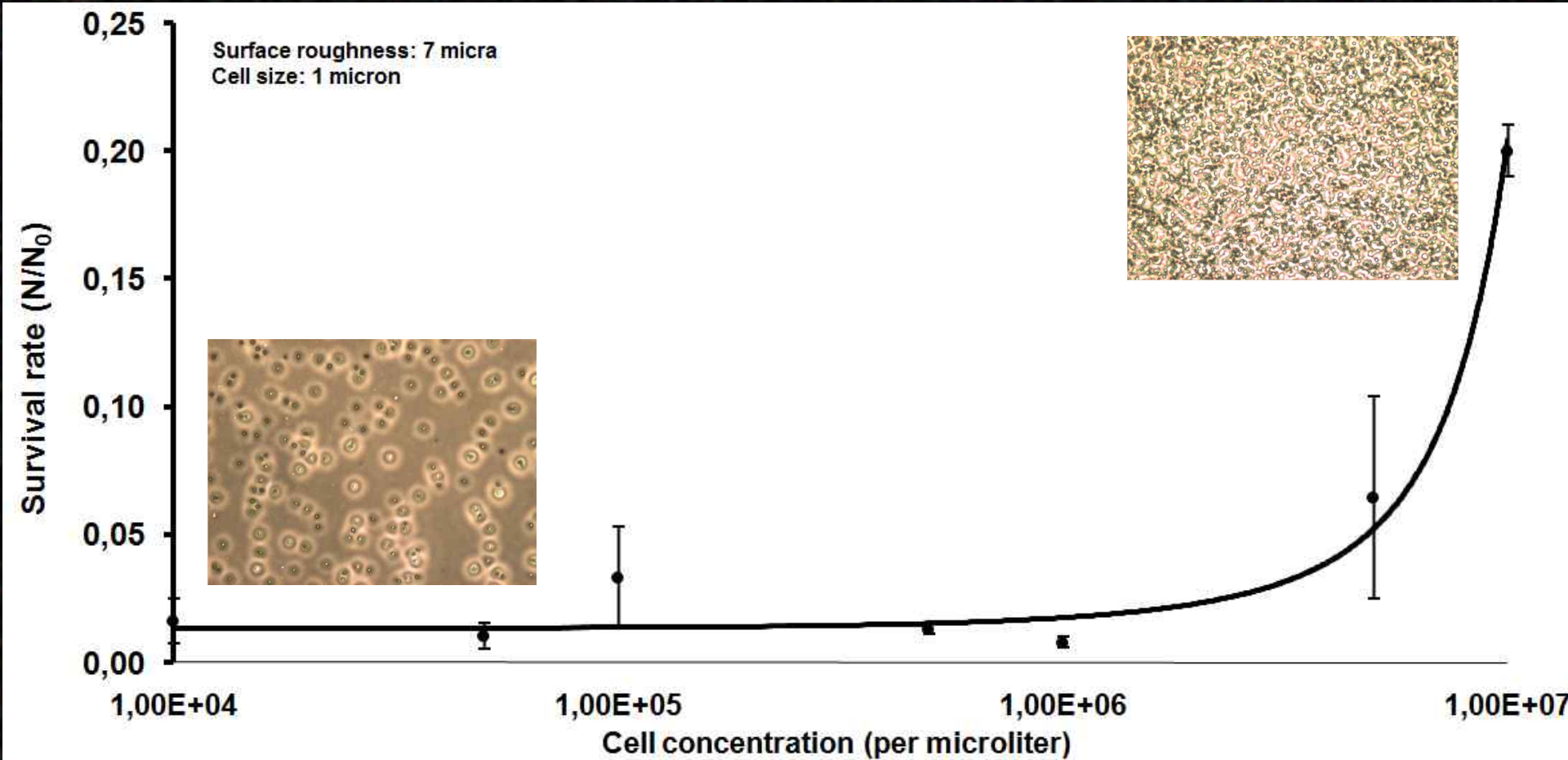


**Microscopic inspection
of carbon tape**

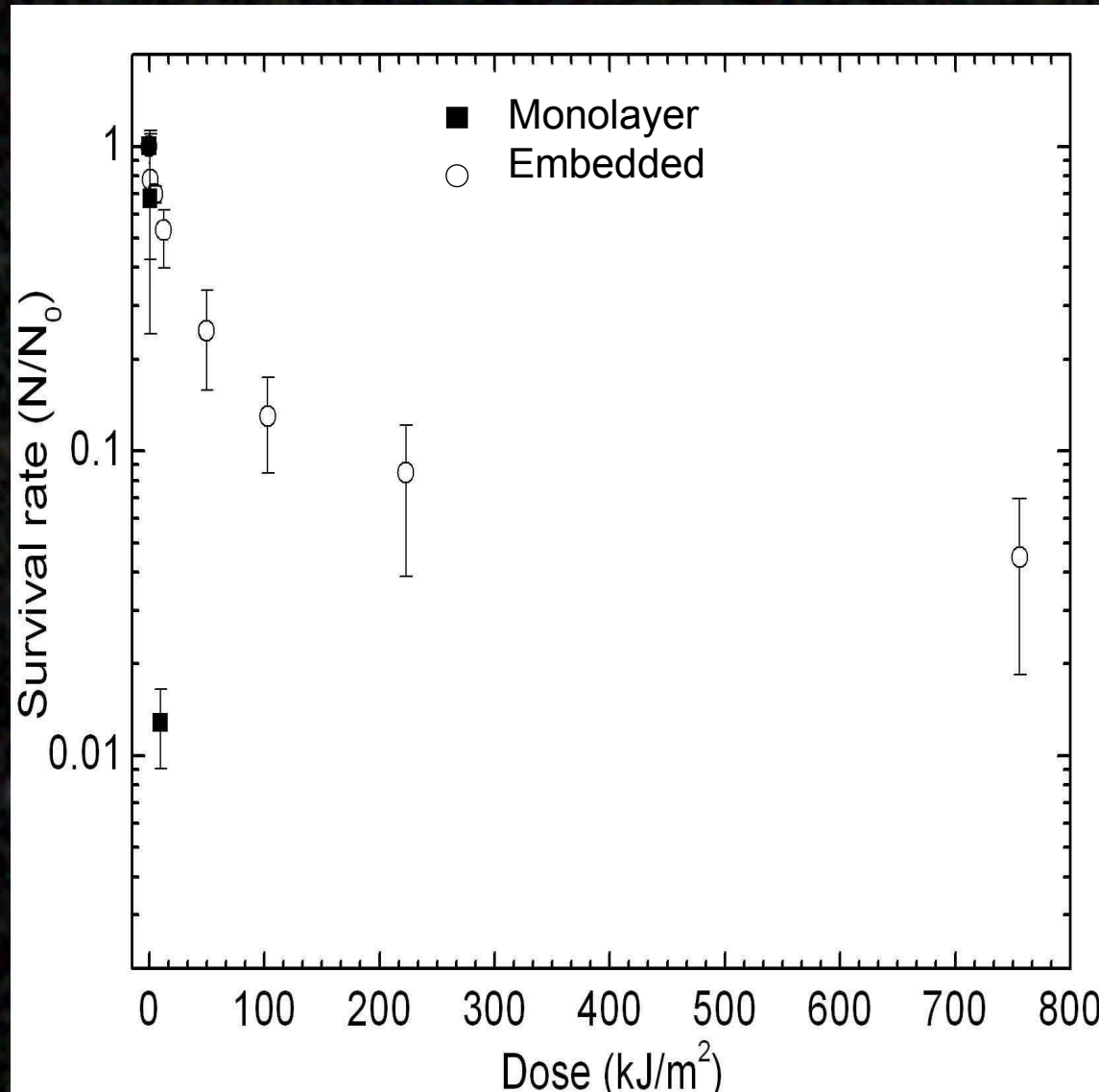


Survival of *D. radiodurans* after exposure to VUV beamlight. Samples were deposited on surfaces with two different levels of roughness (~ 0.7 and $\sim 7 \mu\text{m}$).



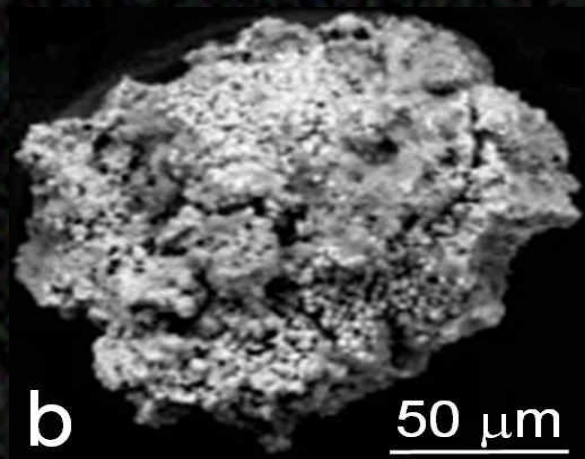


Survival of *D. radiodurans* to irradiation with hydrogen lamp $\lambda=121.6\text{nm}$ (10.2eV)

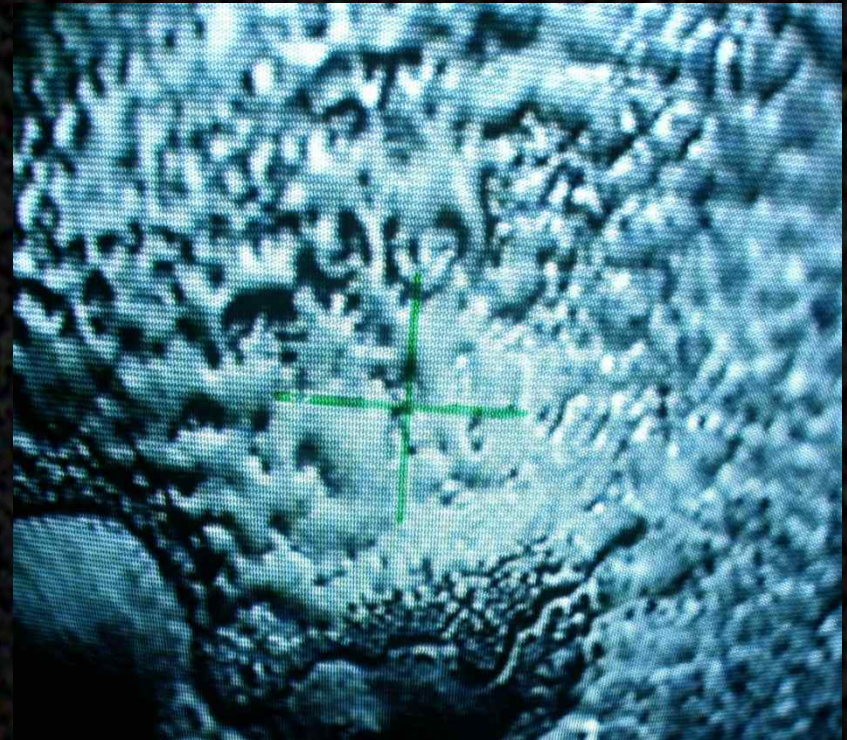


~ 45d of
continuous
solar radiation

CONCORDIA MICROMETEORITES

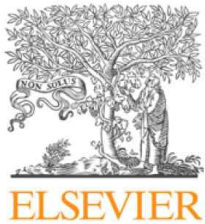


CARBON TAPE



Morphologic comparison between surfaces of Concordia 2002 micrometeorite (Antarctica) and that of the carbon tape on which bacterial powder was layered for irradiation (with permission of M. Maurette).

**ALIENS ARE
MICROBES !!**



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journal homepage: www.elsevier.com/locate/pss



Laboratory simulation of interplanetary ultraviolet radiation (broad spectrum) and its effects on *Deinococcus radiodurans*

Ivan Gláucio Paulino-Lima^{a,*}, Sérgio Pilling^{b,1}, Eduardo Janot-Pacheco^c, Arnaldo Naves de Brito^{b,2}, João Alexandre Ribeiro Gonçalves Barbosa^{b,3}, Alvaro Costa Leitão^a, Claudia de Alencar Santos Lage^a

^a Instituto de Biofísica Carlos Chagas Filho, Avenida Carlos Chagas Filho, 373, Cidade Universitária, CEP 21941-902 Rio de Janeiro, RJ, Brazil

^b Laboratório Nacional de Luz Síncrotron, Rua Giuseppe Máximo Scolfaro, 10000, Pólo II de Alta Tecnologia de Campinas, CEP 13083-970 Campinas, SP, Brazil

^c Instituto de Astronomia, Geofísica e Ciências Atmosféricas, Rua do Matão, 1226, Cidade Universitária, CEP 05508-900 São Paulo, SP, Brazil

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ABSTRACT

The radiation-resistant bacterium *Deinococcus radiodurans* was exposed to a simulated interplanetary UV radiation at the Brazilian Synchrotron Light Laboratory (LNLS). Bacterial samples were irradiated on different substrates to investigate the influence of surface relief on cell survival. The effects of cell multi-layers were also investigated. The ratio of viable microorganisms remained virtually the same (average 2%) for integrated doses from 1.2 to 12 kJ m⁻², corresponding to 16 h of irradiation at most. The asymptotic profiles of the curves, clearly connected to a shielding effect provided by multi-layering cells on a cavitary substrate (carbon tape), means that the inactivation rate may not change significantly along extended periods of exposure to radiation. Such high survival rates reinforce the possibility of an interplanetary transfer of viable microbes.

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